

SPIS-SCI validation studies for fields instruments

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SPIS simulations in support of plasma instruments for Cosmic Vision

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|--------------------|---|
| A. Hilgers (ESA) | A review of spacecraft plasma interactions effects on plasma measurements |
| A. Masson (ESA) | Electron density measurements in the magnetotail with different instruments |
| M. Capacci (Laben) | Observation of SMART-1 plume plasma environment with the EPDP plasma diagnostic package and future activities |
| M. Capacci (Laben) | Charging active control: PLEGPAY experiment onboard ISS results; activities on future systems |
| D. Rodgers (ESA) | Plasma measurements onboard CHAMP spacecraft |
| H. Laakso (ESA) | Observation of spacecraft plasma interactions with Cluster |
| D. Kataria (MSSL) | Spacecraft-plasma interactions: an MSSL perspective |
| A. Eriksson (IRFU) | Cold plasma and electric field measurements in the Jovian system: possibilities and challenges |
| A. Hilgers (ESA) | Modelling of plasma environment of Cluster electrostatic sensors |
| A. Eriksson (IRFU) | Wakes in cold tenuous plasmas: nuisance and blessing |
| S. Clucas (ESA) | MMS electrostatic environment simulation |
| D. Rodgers (ESA) | Champ and Swarm plasma environment modelling |

E-field instruments & Langmuir probes

- Rely on electric coupling to plasma, sensitive to s/c-plasma issues
- Asymmetric antenna or s/c configuration
- Photoelectron clouds and currents
- Secondary emission currents
- Wake potential, asymmetric shielding

Example missions

- Some missions needing instrument simulations:
 - In space: Rosetta, Cassini, Cluster, THEMIS
 - Upcoming: Swarm, BepiColombo, MMS
 - Design phase: JUICE, Solar Orbiter, SP+
- In most cases, wide ranges of plasma parameters are encountered
 - Example: Debye lengths for Rosetta vary from a fraction of a mm (fully developed inner coma) to tens of meters (tenuous solar wind at 3 AU)
 - No single simulation setup can cover all this with just a change of parameter values
- Some include thin (mm) and long (tens of m) wire booms challenging to model
 - Cluster, THEMIS, MMS, BepiColombo MMO

SPIS-SCI developments

- Many important improvements for science users, for example:
 - Backtracking. Current to small areas (sensors) on big s/c can be measured at good accuracy without excessive number of particles.
 - Particle instruments and Langmuir probes. Possible to define detailed geometry, backtrack currents and step bias potential.
 - Photoemitting thin wires. Enables realistic modeling of influence on probes from nearby elements.
 - Parallellized particle pushing.

SPIS-SCI validation aims (field instruments)

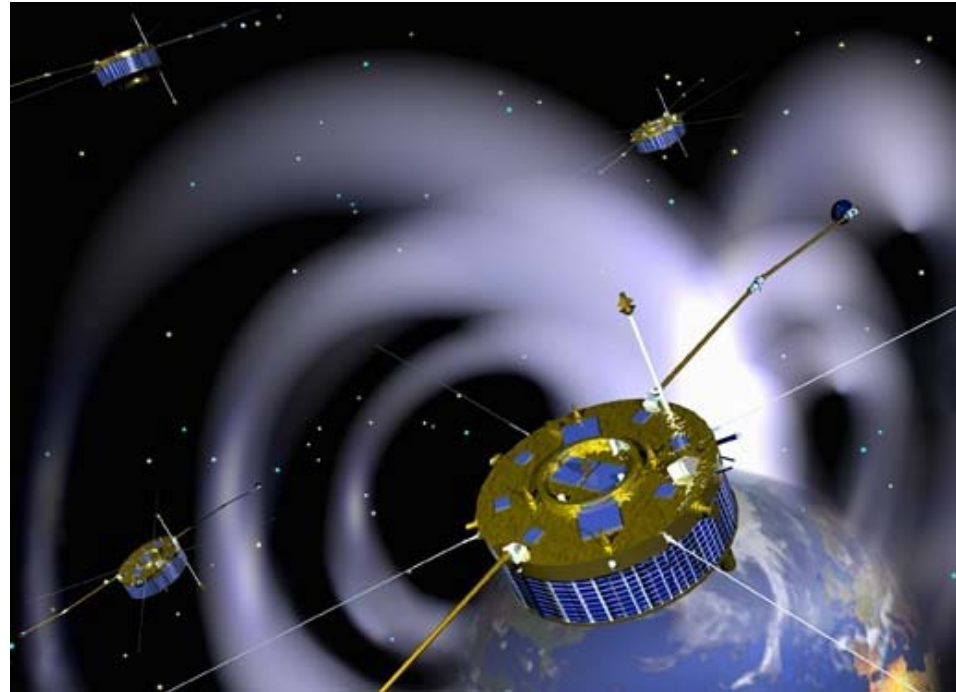
- Verify SPIS-SCI can be used for studying performance of E-field instruments and Langmuir probes
- Comparisons where possible:
 - Simplified analytical expressions
 - Other simulation results
 - Data

Cluster, THEMIS, MMS, Bepi MMO

- Missions with long wire boom electric fields
- Problematic to simulate in old SPIS
 - Particularly including photoemission
- SPIS simulations by Prakash (2007)
- Problem:
 - Large scale size disparity (mm to hundreds of m)
 - Simulation box needs to be hundreds of meters
 - Enormous number of particles needed in strict PIC approach to get reasonable number of particles to a small probe
- SPIS-SCI: Backtracking solves scale disparity issue
- SPIS-SCI also allows photoemitting thin wires, which also can be used as Langmuir probes

Cluster wire boom E-field instrument

- Cluster: 4 ESA s/c in orbit since 2000
- $\approx 2 \times 20$ RE
- Wire booms 88 m tip2tip



Cluster E-field features to study

- Sunward offset
 - Always in double probe instruments in tenuous plasmas
 - Around 1 mV/m on Cluster
 - Attributed to asymmetric photoemission
 - First simulated by Cully et al (2007)
- Enhanced wake
 - Ion flow diverted by strongly positive s/c in tenuous plasma
 - Can give signatures of up to 10 mV/m in data
 - First simulated by Engwall et al (2006)
 - No previous simulation includes the probes themselves

Biased elements

- Fig from Cully et al, JGR 2007
- Potentials:
 - Probe at V_p , about 1 V positive to local plasma
 - Thin wire at V_p
 - Guard at $V_p - 6$ V
 - Puck at $V_p + 1.3$ V
 - Boom at V_{sc}

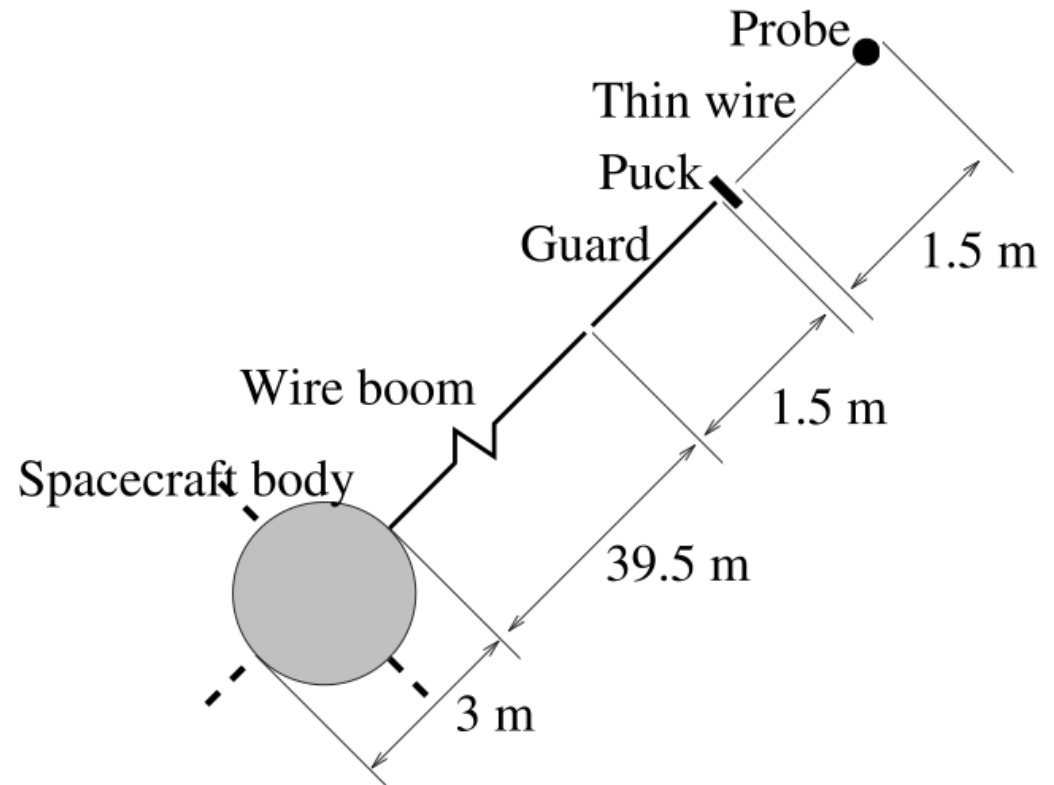
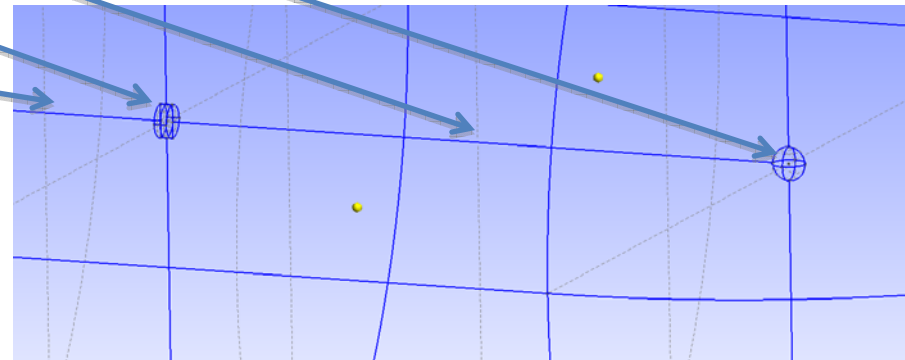
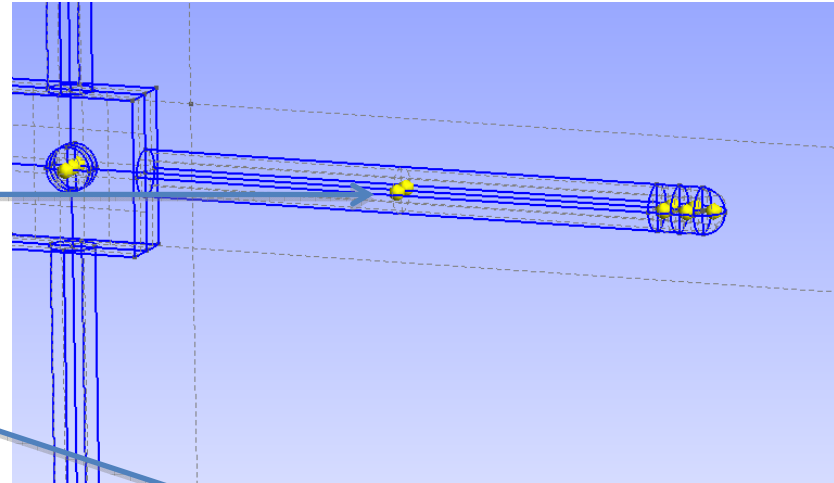
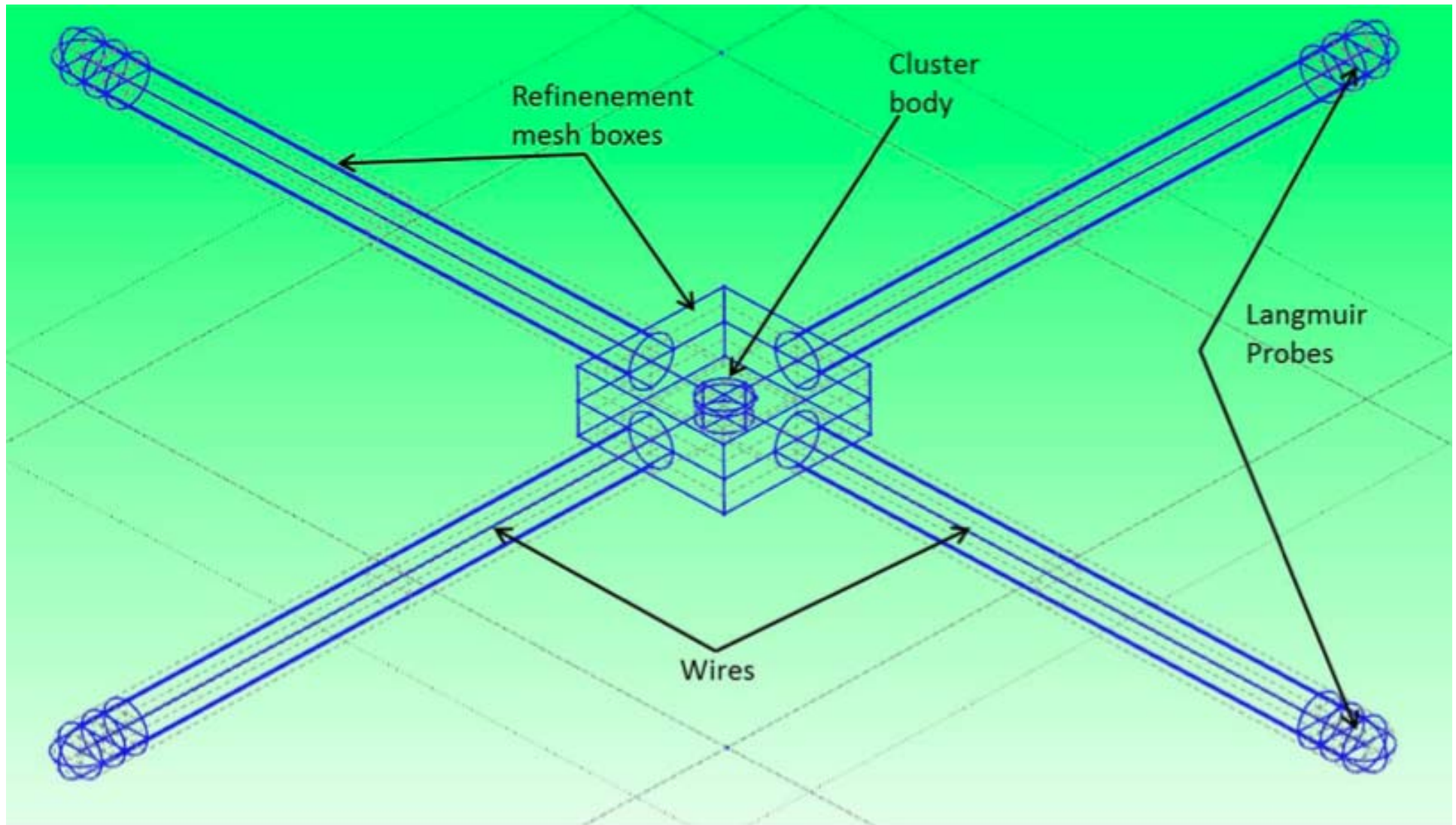


Figure 1. Geometry of the Cluster spacecraft and the EFW wire booms. There are four distinct electrical elements: (1) the spacecraft and wire booms, at a potential determined by the current balance in the plasma (but specified as a parameter in the simulation); (2) the guard, normally biased to 6 V below the probe; (3) the roughly cylindrical puck (preamplifier enclosure) (diameter of 8 cm and length of 3 cm), normally biased to 1 V above the probe; and (4) the thin (0.3 mm radius) wire and the probe, connected with a constant bias current to the spacecraft. The probe is a sphere with a diameter of 8 cm, and the wire radius (including the guard) is 1.1 mm.

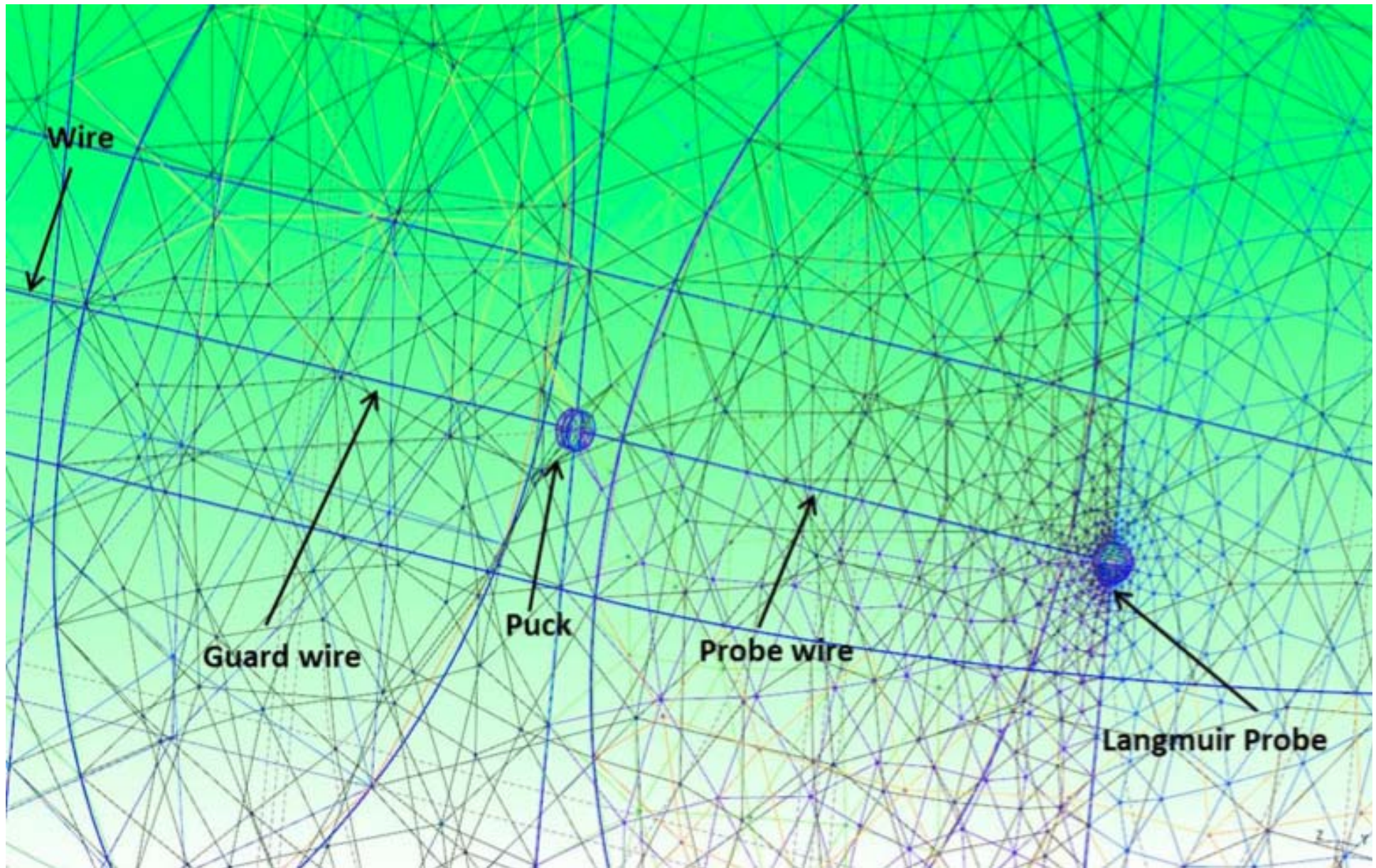
SPIS CAD model

- Biased elements
 - boom (V_{sc})
 - sphere (V_p)
 - thin wire (V_p)
 - puck ($V_p + 1.3 \text{ V}$)
 - guard ($V_p - 6 \text{ V}$)
- Need to simulate all four probes
- Very demanding simulation case



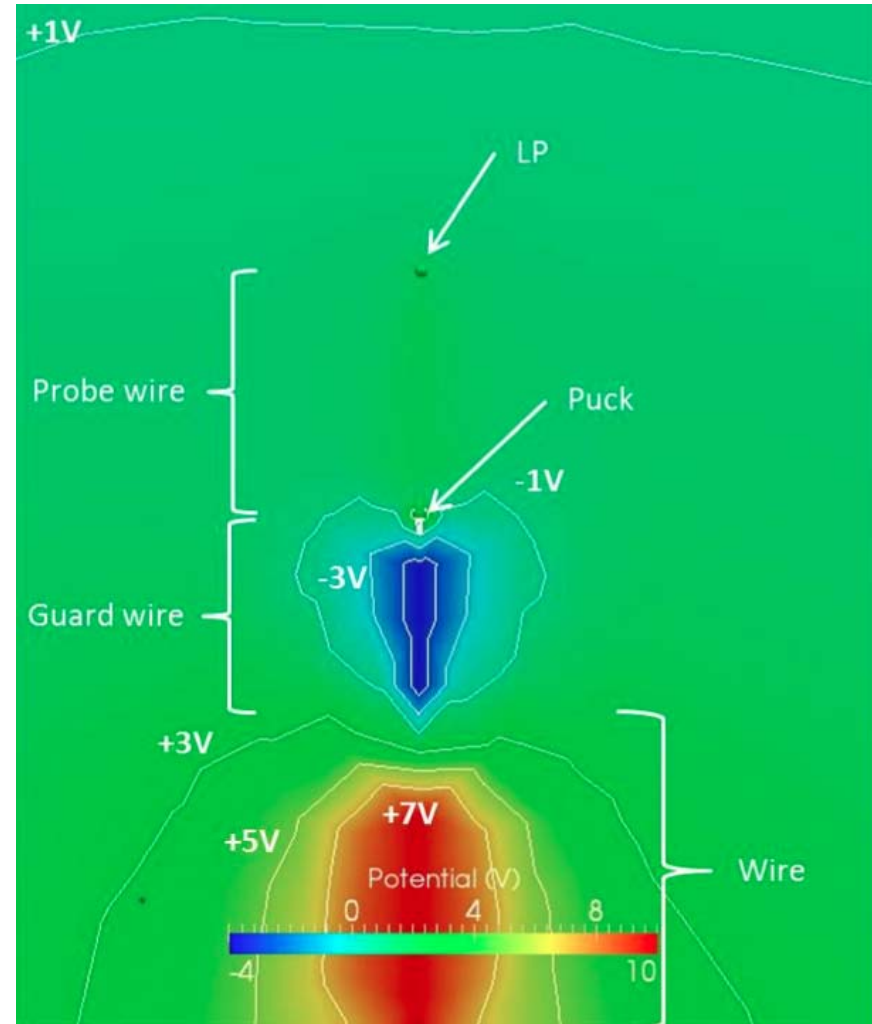
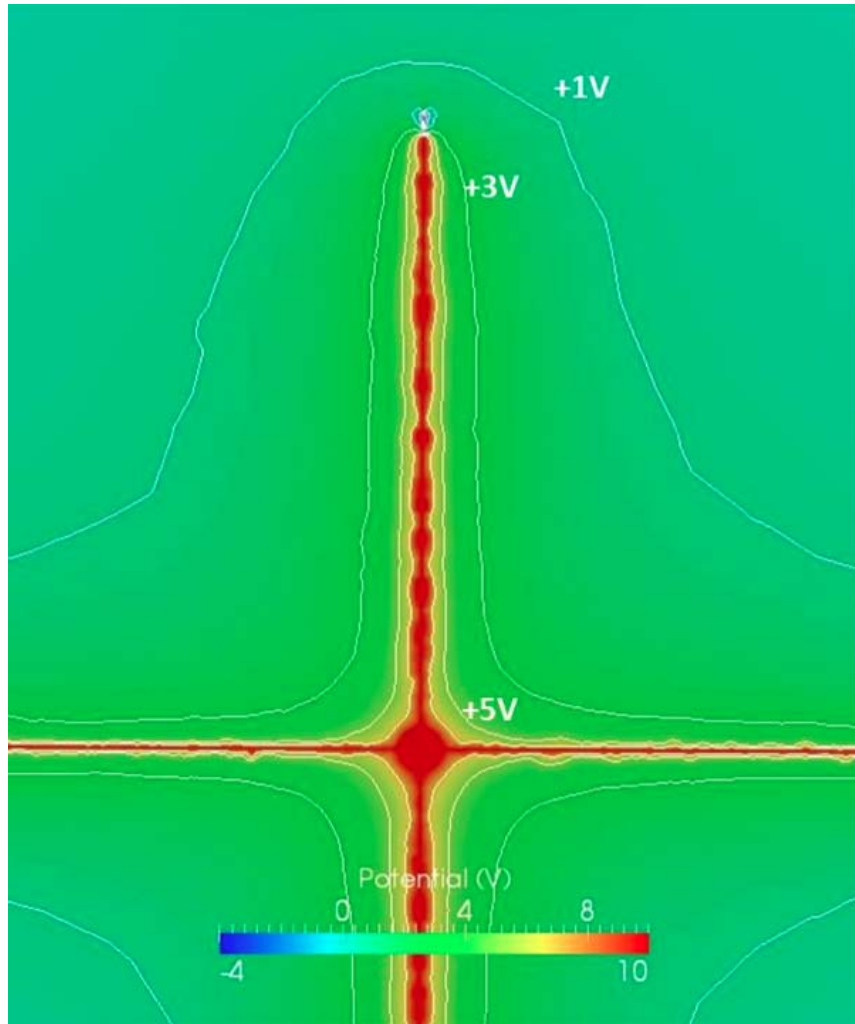


Cluster geometric model: large scale



Cluster geometric model: details of mesh around probe

Cluster in solar wind ($V_{sc} 10 \text{ V}$)

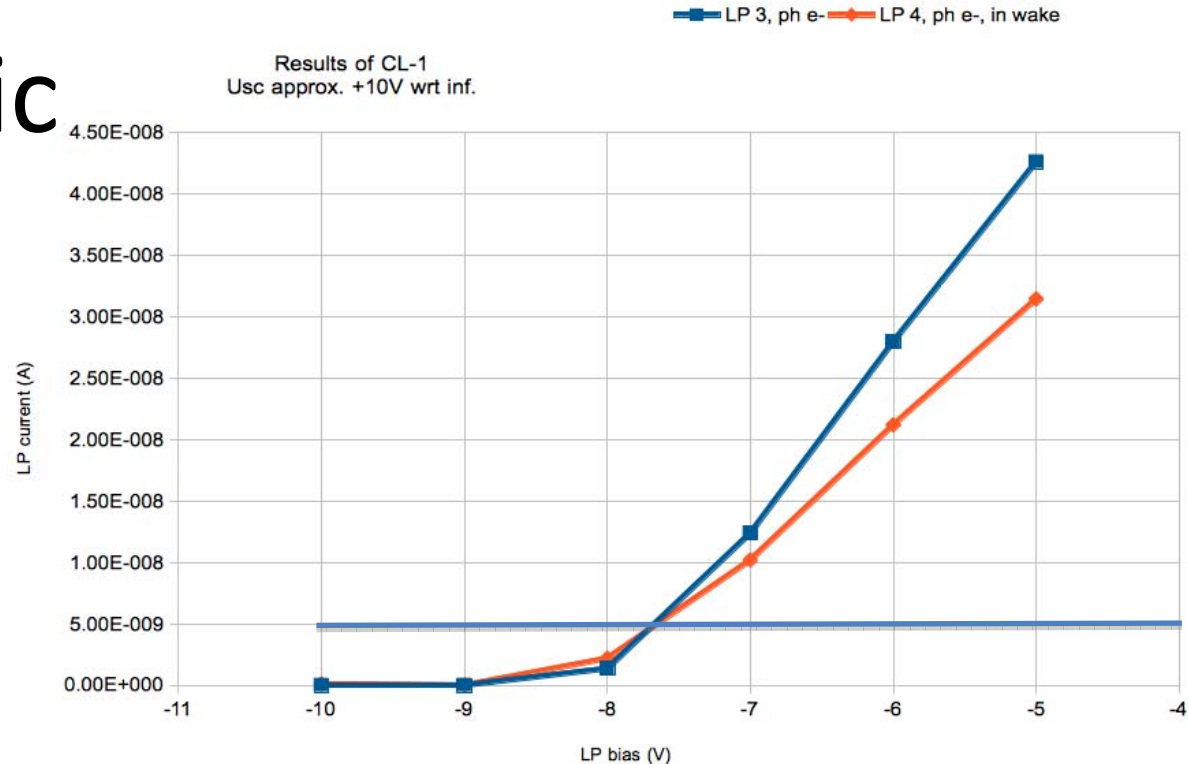


SPIS-SCI probe characteristic

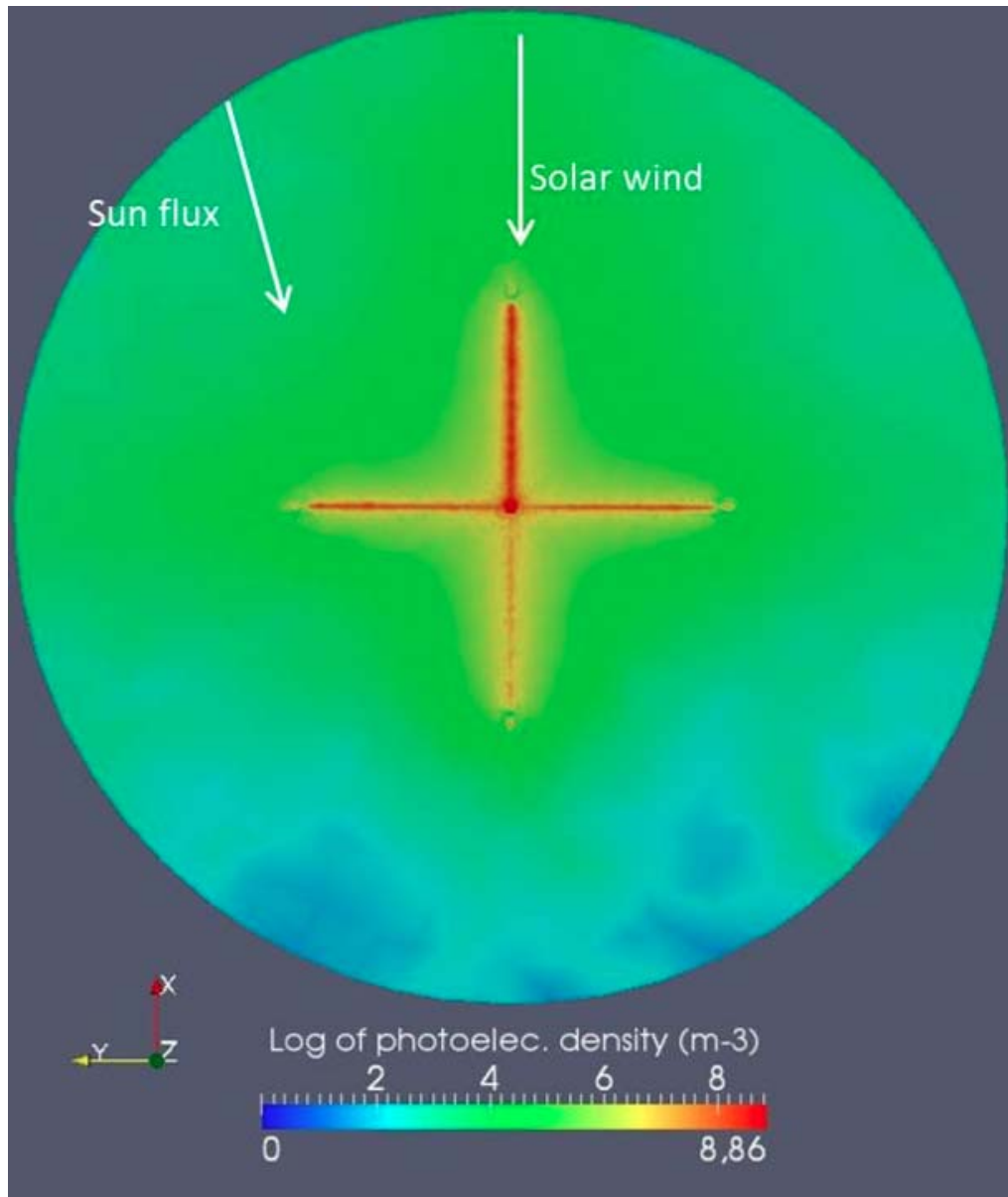
- Simulated for v opposing probe (sunward-antisunward)
- Tenuous solar v plasma
- Horizontal shift at given (bias) current gives sunward offset

Photo e-

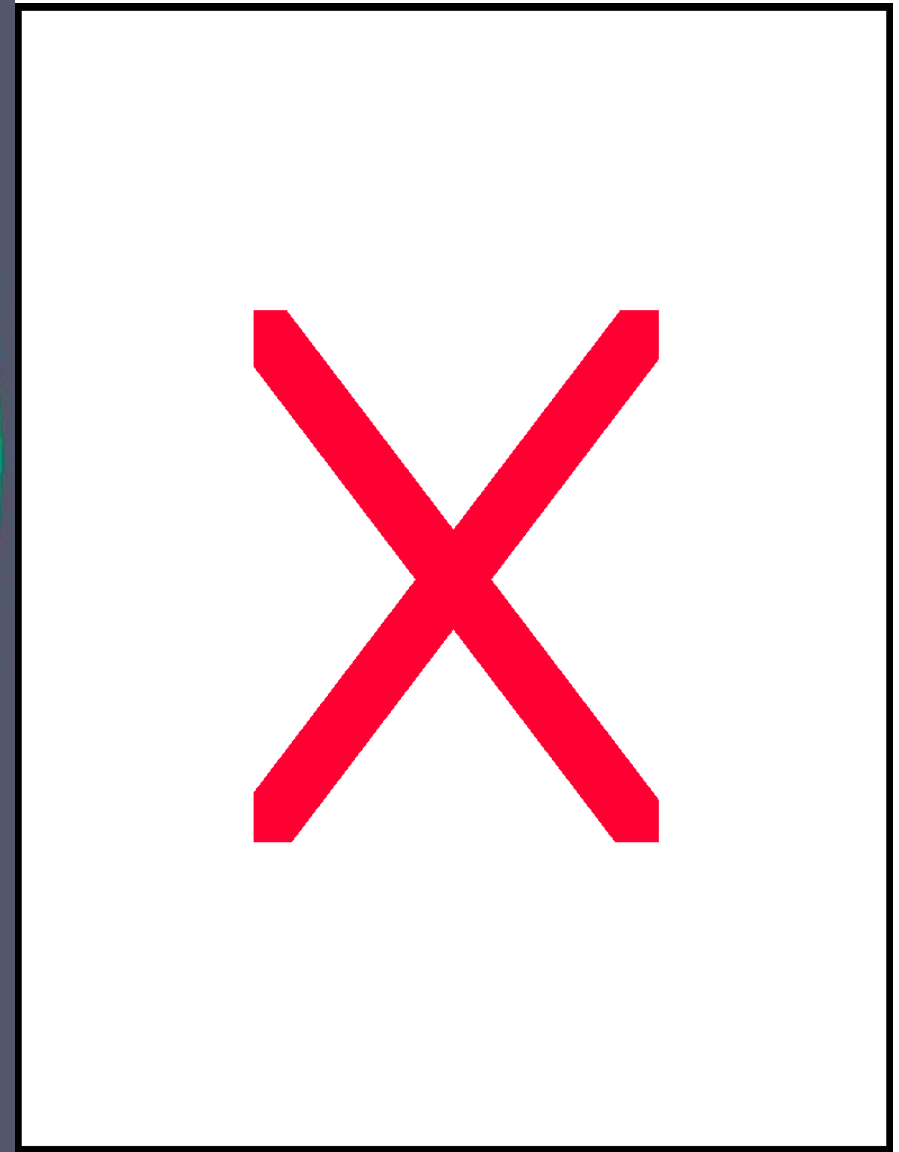
Results of CL-1
Usc approx. +10V wrt inf.



- Possible to study performance as function of environment, spin phase, UV flux...
 - Will require lots of CPU time, but is now possible



Guillemant/Nilsson SPIS-SCI simulation
($V_{sc} = 10$ V)



Cully et al (2007) ($V_{sc} = 25$ V)

What does this mean?

- Could be done before:
 - SPIS: Simulate potential pattern neglecting influence of thin wire photoemission
 - Cully vacuum code: Instrument performance in long Debye lengths
 - Miyake et al PIC code: Instrument performance in any plasma for distorted instrument geometry
- Can be done in SPIS-SCI:
 - Potential pattern including photoemission and secondary emission from every element
 - Instrument performance in any plasma
 - Detailed and realistic geometric model

Conclusions

- SPIS-SCI can be used to evaluate performance of plasma and field instruments
 - Particle trajectories
 - Blocking by s/c
 - Photoemission, secondaries
- The software demands an effort by the user
 - Need for training events
 - Need for support and documentation
 - Also a community issue – we can all help
- Parallelization of particle pusher means simulations take days rather than weeks
- SPIS-SCI really gives new possibilities in this field!