

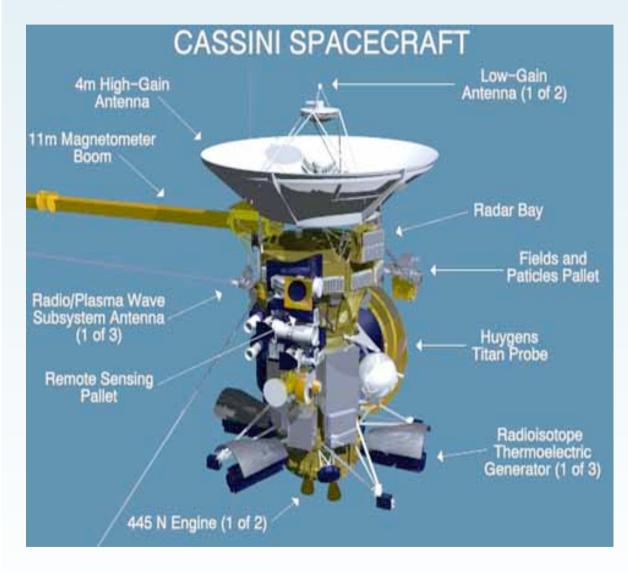
# A study of photoelectrons on the Cassini spacecraft using the SPINE software

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#### Cassini: an overview



- ➤ The most advanced spacecraft to be launched
- ➤ 12 instruments on board
- Complex shape. Lots of irregular metallic surfaces
- > Tri axis stabilised
- Operates in different plasma regimes



## **ELS:** an overview

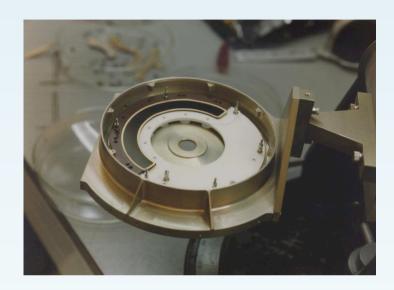


- Electron Spectrometer (ELS)
- ➤ Ion Mass Spectrometer (IMS)
- ➤ Ion Beam Spectrometer (IBS)
- Collectively form the Cassini Plasma Spectrometer (CAPS)





## **ELS:** an overview



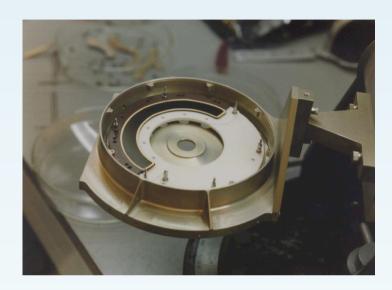
- ➤ Various parts of the spacecraft affect the electron measurements
- ➤ The most 'clear view' is at Anode 4, Act=0

- ➤ Inside the ELS, electrons are collected over 8 anodes that cover 160°
- ➤ The ELS is mounted on an 'Actuator' that rotates through ~200°



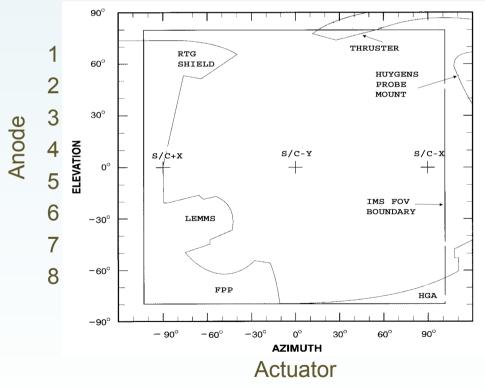


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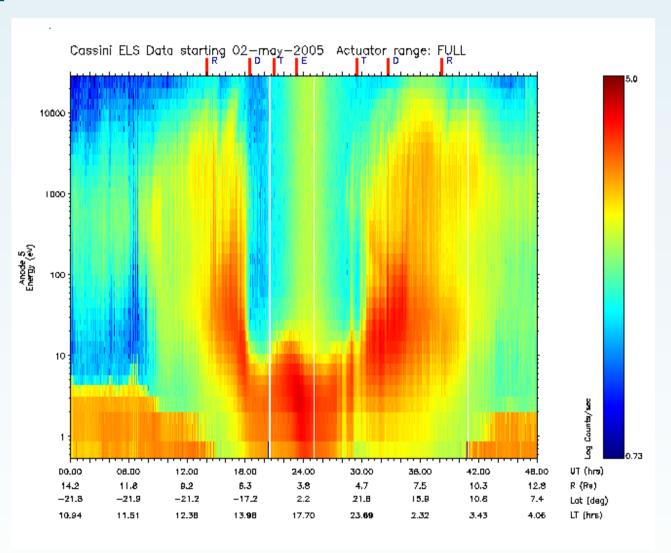
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#### **ELS:** the data

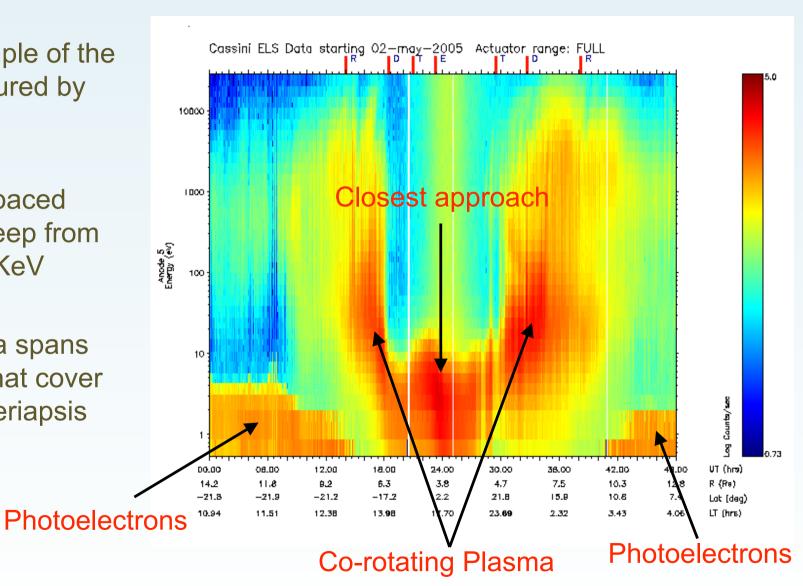
- ➤ An example of the data measured by the ELS
- ➤ 63 log spaced energy sweep from 0.5eV 26KeV
- ➤ This data spans two days that cover a Saturn periapsis





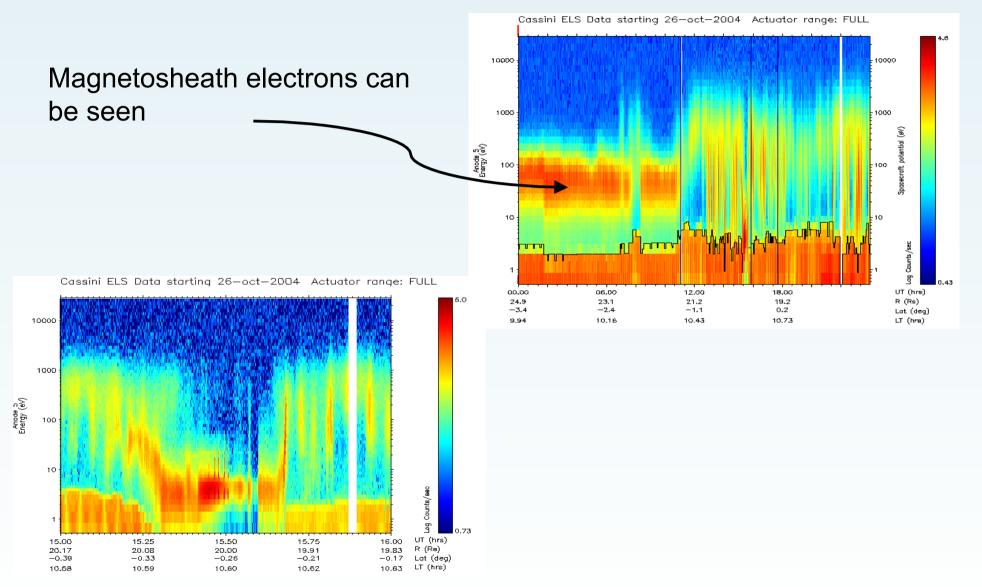
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## A wealth of information

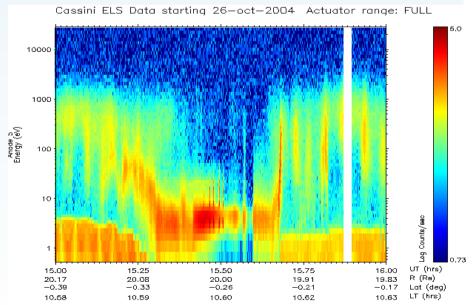




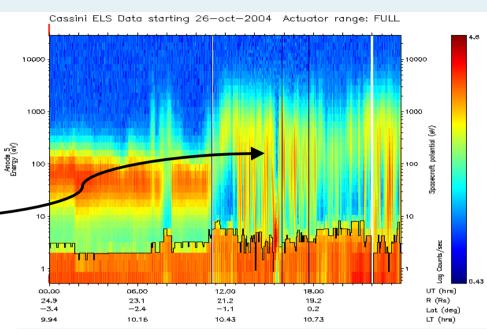
## A wealth of information

Magnetosheath electrons can be seen

and outer magnetspheric electrons



15th SPINE meeting, ONERA, Toulouse

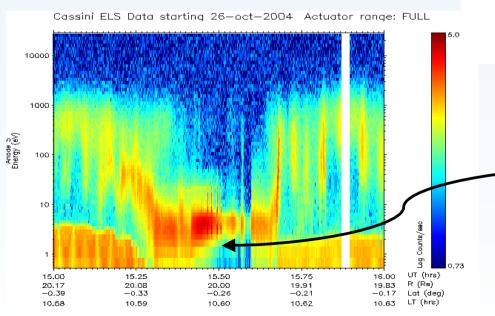


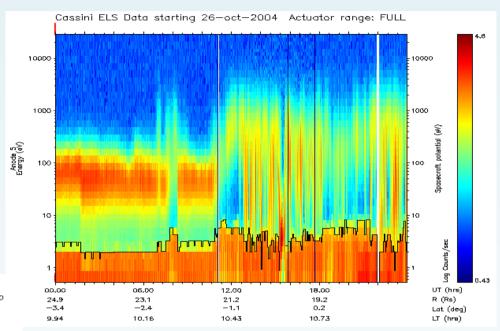


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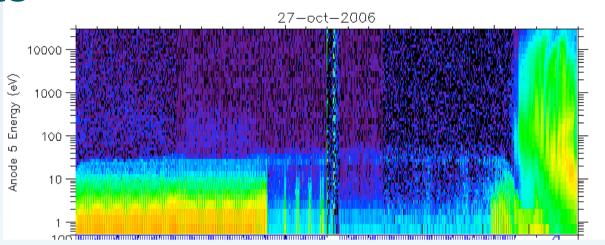
Here photoelectrons disappear during the Titan encounters due to high density

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28th-29th September 2009

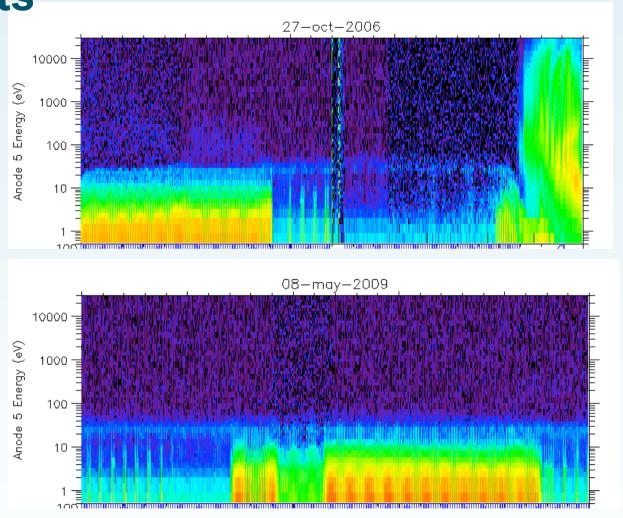


- An example of the variation in photoelectron distribution
- SC close to periapsis
- ➤ Are these ionospheric photoelectrons?



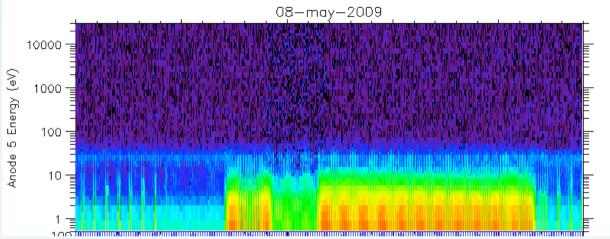


- ➤ An example of the variation in photoelectron distribution
- SC close to periapsis
- ➤ Are these ionospheric photoelectrons?
- > No.



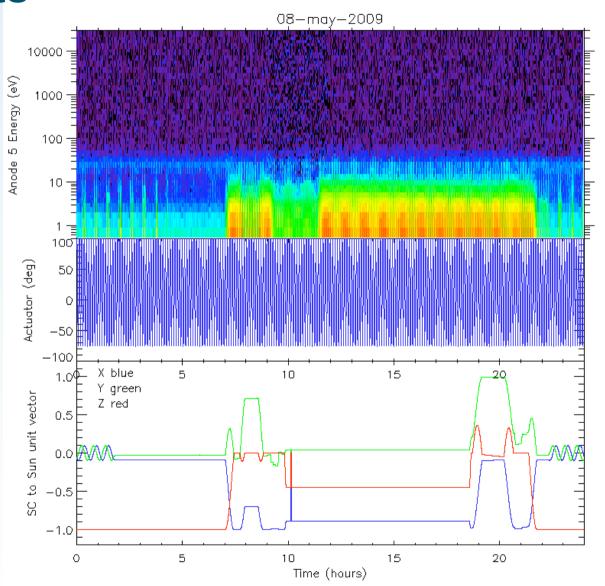


So what's causing these effect?





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- Unit vector of SC to Sun
- ➤ HGA points to Sun

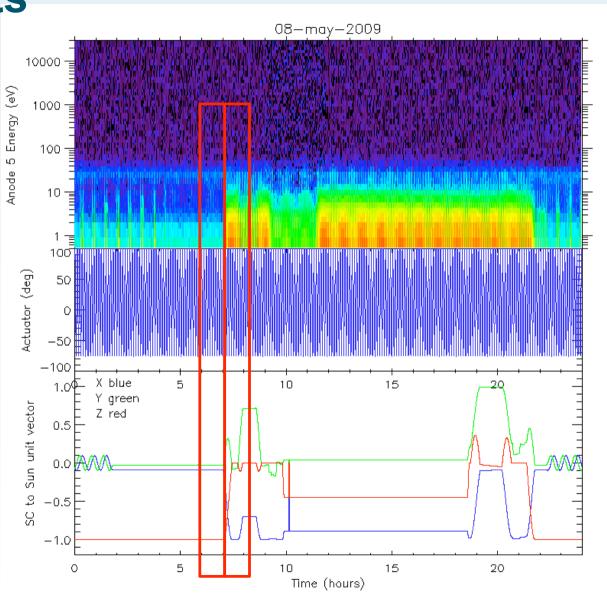




- So what's causing these effect?
- Unit vector of SC to Sun
- > HGA points to Sun
- ➤ Average data from

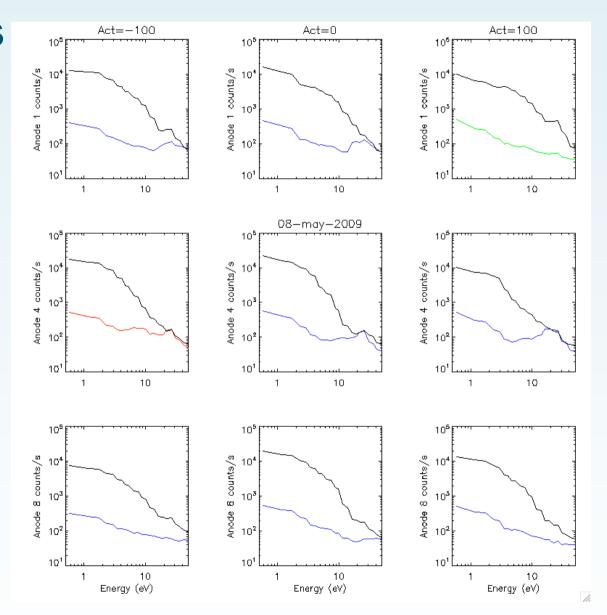
♦ A: 06:00-07:00

♦ B: 07:00-08:00



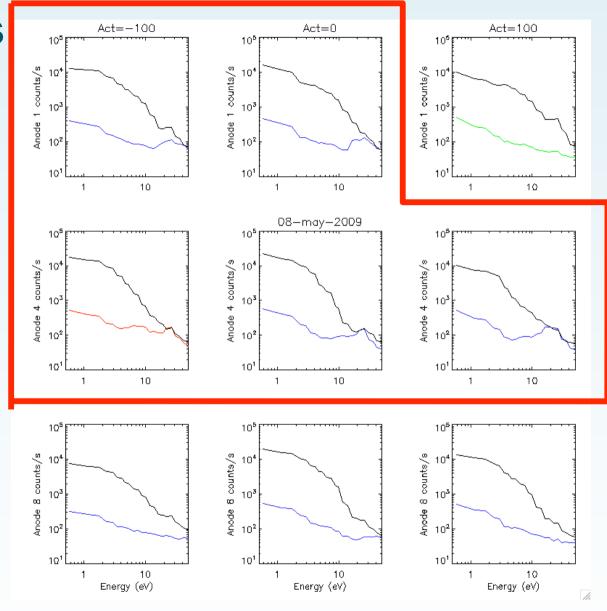


- ➤ Black lines from Data B
- Coloured lines from Data A





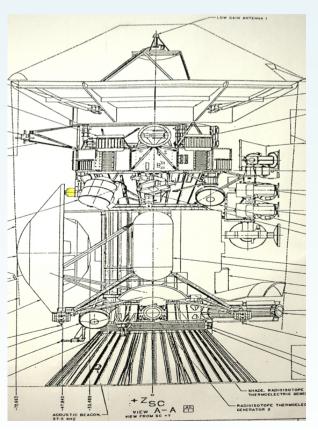
- ➤ Black lines from Data B
- Coloured lines from Data A
- ≥ 25 eV peak is dependent on direction

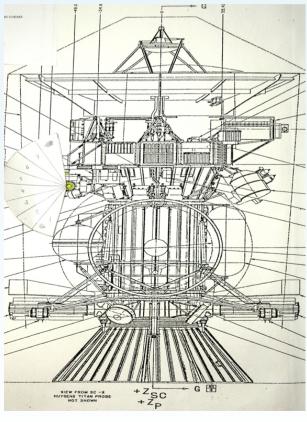


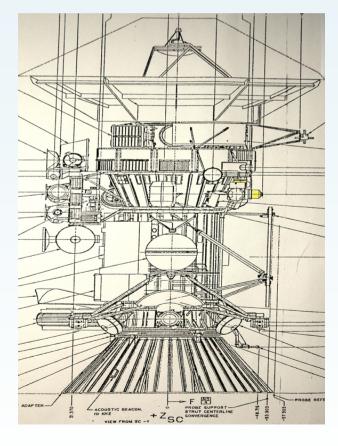


## **SC Layout**

➤ The space craft diagrams show a possible origin of the photoelectrons based on anode.....





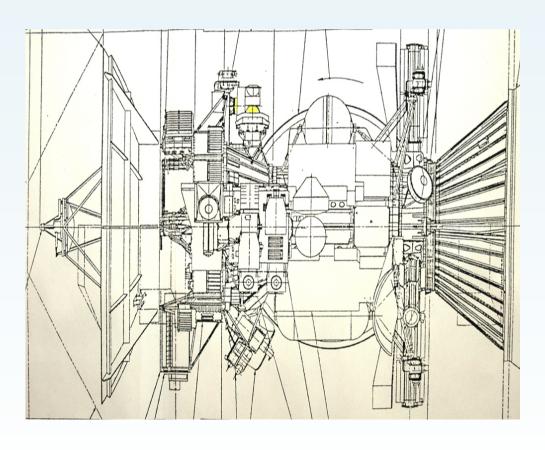


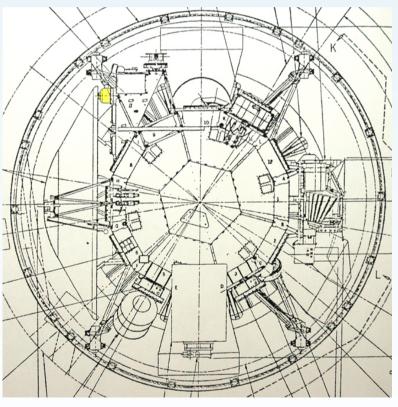
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## **SC Layout**

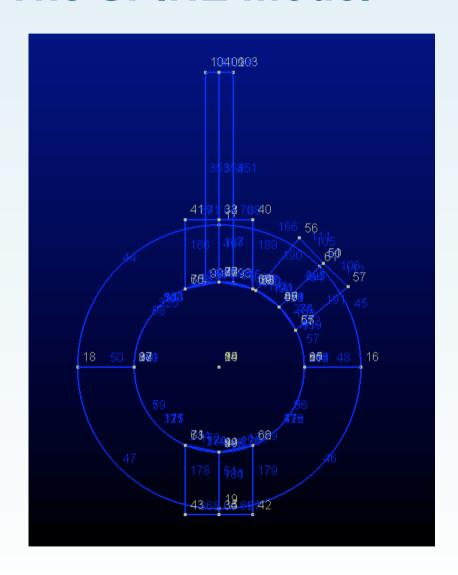
- ➤ The space craft diagrams show a possible origin of the photoelectrons based on anode......
- > .....and actuator

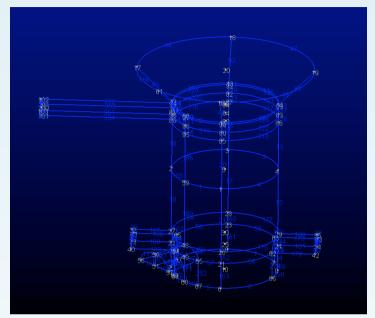


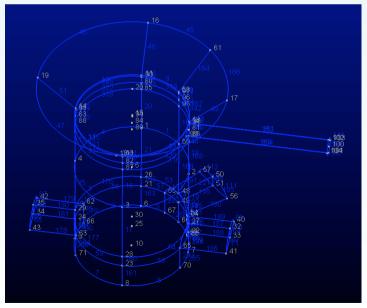




## The SPINE model







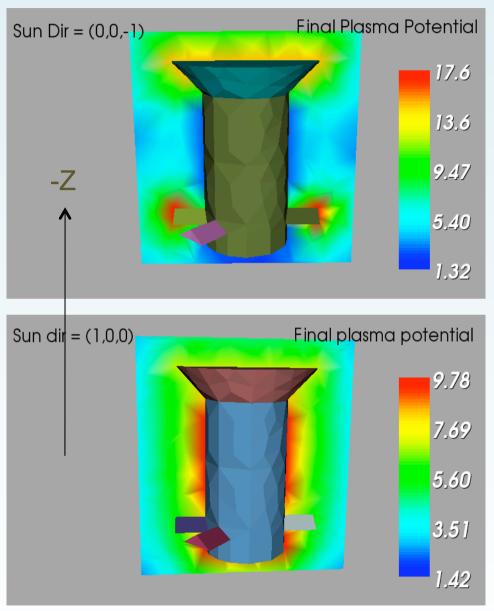
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#### **SPINE** results

- HGA-steel
- RTG shades-teflon
- SC body-gold
- Initial electron density=800 m<sup>-3</sup>
- > Low potential in shadow regions
- ➤ High potential on tips of RTG shades
- ➤ High potential on sc body





## **Summary**

- Beam at 25eV may be coming from RTG shades
  - ♦ SPINE is very useful in studying the photoelectron distributions
  - ♦ Much more work to be done
  - ♦ Extend the model.
  - ♦ Actual materials
  - ♦ Much to learn about SPINE
- Future plans
  - ♦ More advanced model
  - ♦ Different magnetospheric region comparisons
  - ♦ Particle trajectories?
- Further uses of SPINE at MSSL
  - Cassini (negative SC charging problem)
  - ♦ VEX (Solar Panels)
  - Europa Jupiter System Mission (EJSM)



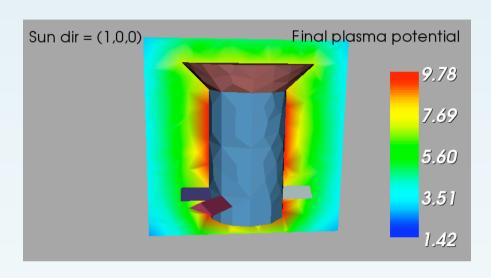
grl@mssl.ucl.ac.uk

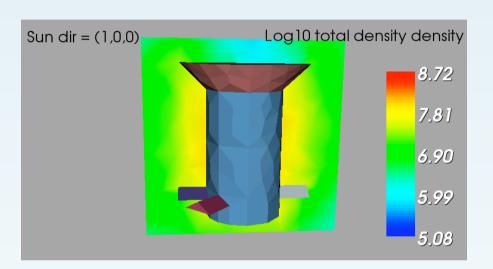


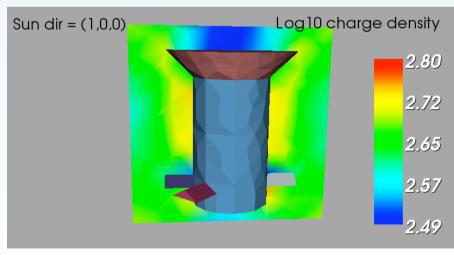
#### > Extras

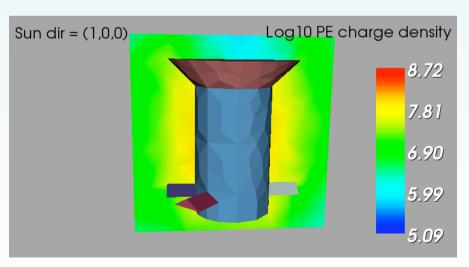


#### **SPINE** results











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