



SPIS status

SPINE Workshop, March 7, 2012



retour sur innovation

SPIS context and project overview

- SPINE (Spacecraft Plasma Interaction Network in Europe) community setup around year 2000 (A. Hilgers, J. Forest, JF Roussel...):
 - An idea was born: gather European efforts for SC-plasma interactions
 - Exchange: knowledge, data, codes, results...
 - Boost the development of a common simulation toolkit: ESA ITT in 2002 => SPIS
- SPIS Development (Spacecraft Plasma Interaction Software) :
 - Initial development: 2002 – 2005
 - ONERA-Artenum consortium
 - ESA/ESTEC TRP contract
- Solver enhancement: 2006 – 2009
 - Mostly ONERA
 - ESTEC ARTES contract, French funding
- Other
 - Some community developments
 - Some CNES-funded enhancements (EP, ESD)
 - ESD triggering modelling (ESA TRP) 2010
- 2011 : SPIS version 4.3.1

SPIS on-going activities

- at ONERA and ARTENUM
 - SPIS-SCIENCE
 - Development of a "Computational tools for spacecraft electrostatic cleanliness and payload accommodation analysis"
 - Next slides
 - SPIS DC-Deep Charging
 - Development of an Internal Charging solver in the frame of the ELSHIELD project
 - Presentation by P. Sarrailh (ONERA)
 - SPIS-GEO
 - Development of a "Simplified Standard MEO/GEO Tools for Spacecraft Charging"
 - Presentation by J. Forest (ARTENUM)
 - SPIS-Maintenance (ESA TRP)
 - "Winter School" on SPIS use
 - in conjunction with this SPINE WS
 - 2 training sessions of 10-12 students (1st one was yesterday; next is tomorrow !)
 - Definition of non regression procedure : ~10 "basic" spis test cases provided to the community
- And also
 - SPIS-PROPU
 - Improvements in SPIS in the frame of AISEPS
 - Presentation by M.Wartelski (Astrium F)



SPIS-SCIENCE progress

SPINE Workshop, March 7, 2012

ONERA :	J.-C. Matéo-Vélez, P. Sarrailh, J.-F. Roussel
ARTENUM :	J. Forest, B. Thiébault, B. Ruard
IRF :	A. Eriksson
IRAP :	V. Génot
ESA:	A. Hilgers



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Outline

- Overview of the activity
 - Scope
 - Objectives
- Work progress
 - User requirement from scientific community for missions modeling
 - SPIS-SCIENCE tool
 - Overview of main features
 - Some details on progress and work to be performed
 - Validation campaign
- Summary

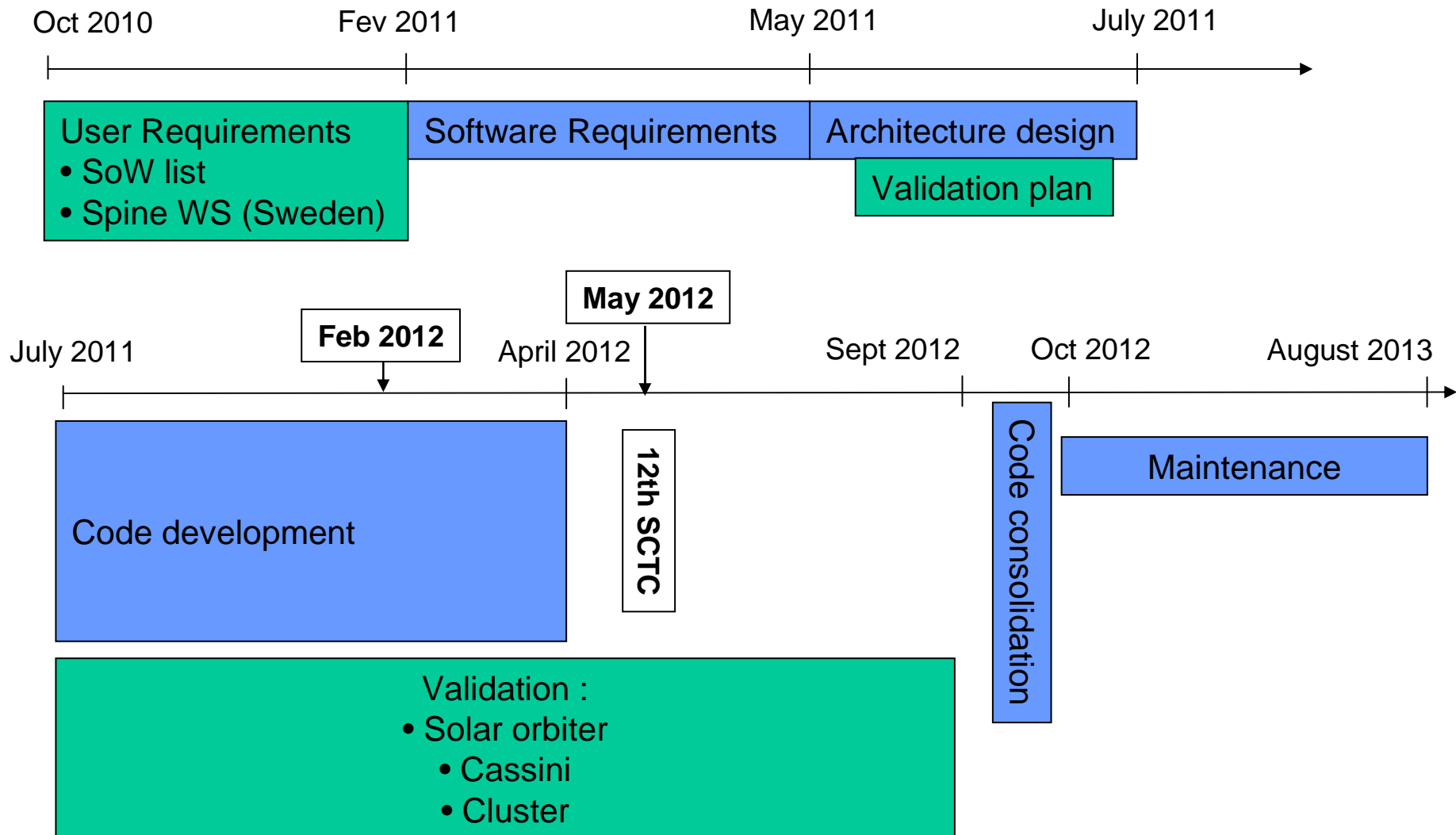
Overview of the activity

- Scope
 - ESA/ESTEC contract 4000102091/10/NL/AS
 - Technical Officer: Alain Hilgers
 - Partners: ONERA, ARTENUM, IRAP, IRFU
 - Long-term scientific program of ESA has planned missions dealing with plasma measurements
 - Solar Orbiter, Juice (Jupiter)
 - Relatively low energy (few eV) plasma measurements
 - Electrostatic cleanliness becomes very important
- Objectives
 - Provide a computational tool able to predict quantitatively
 - The charging of a S/C
 - The space charge in its environment
 - Their consequences in low energy plasma measurements
 - Outputs for the user (= scientific community)
 - New version of SPIS software
 - Validation test cases related to the scientific mission conditions
 - One year maintenance

Work progress

ONERA/ARTENUM

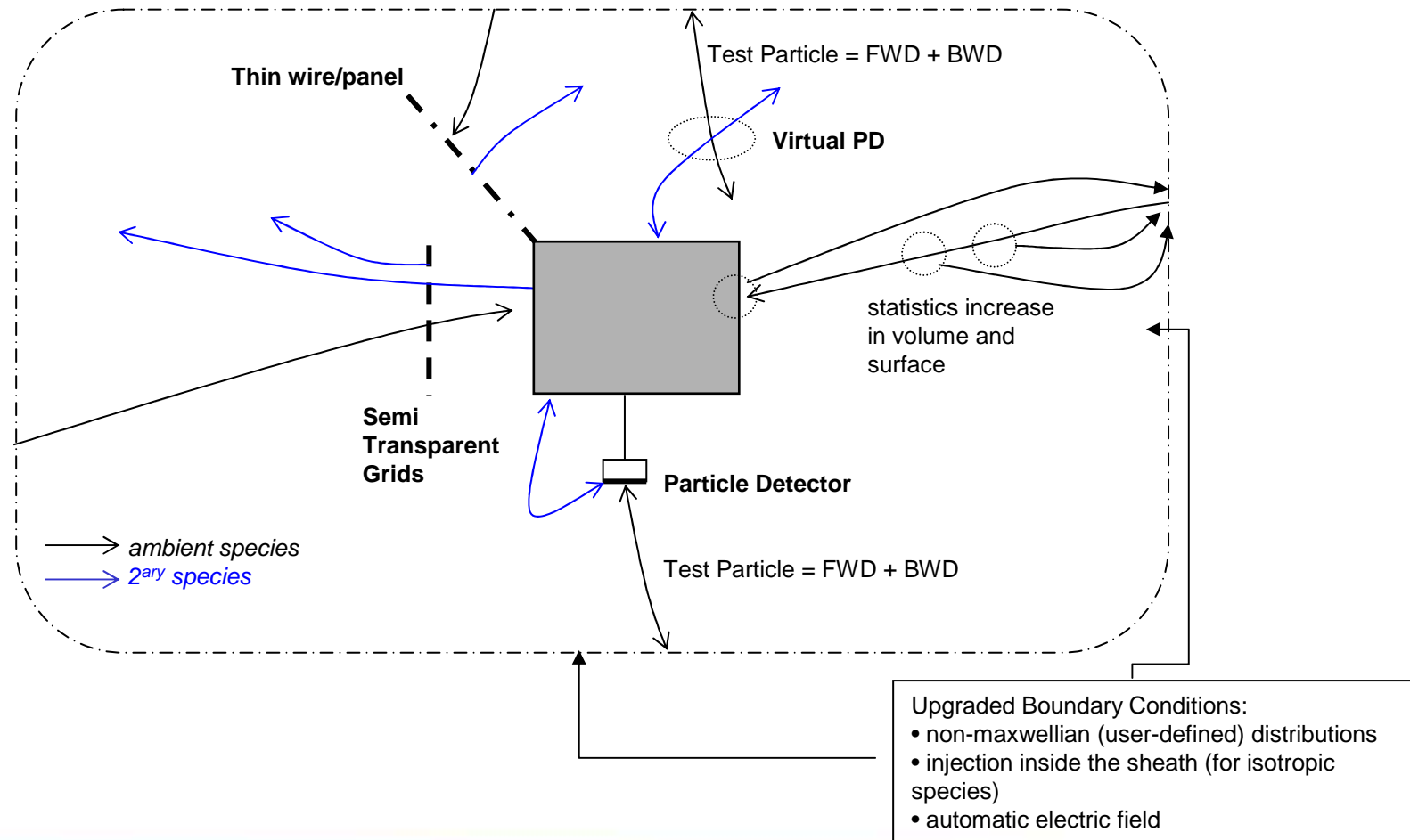
IRAP/IRFU



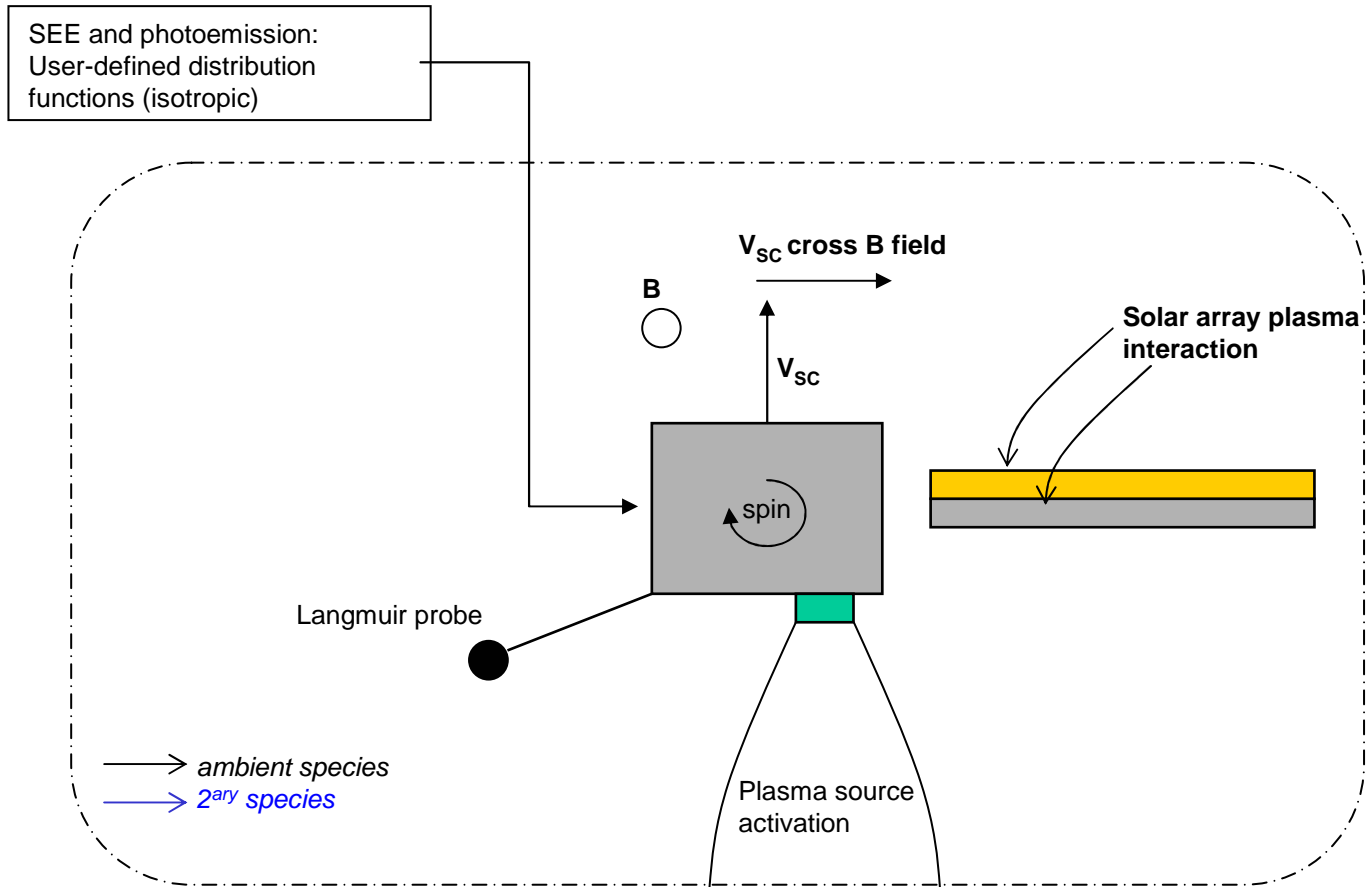
Scientific missions modeling – User requirements

- User requirements for an upgraded SPIS
 - From past experience of ESA and consortium
 - Organization of a SPINE WS at Uppsala, Jan 2011, to collect UR from the relevant science community
 - ~ 50 attendees from scientific univ. or instit., space industry, agencies
 - ~ 25 presentations
 - Plasma physics modeling, particle and plasma instruments, flight measurements, mission needs, SPIS use and limitations ...
 - Exhaustive and not limited user requirements list
 - Detailed requirements on phenomena modeling

Work in progress



Work in progress



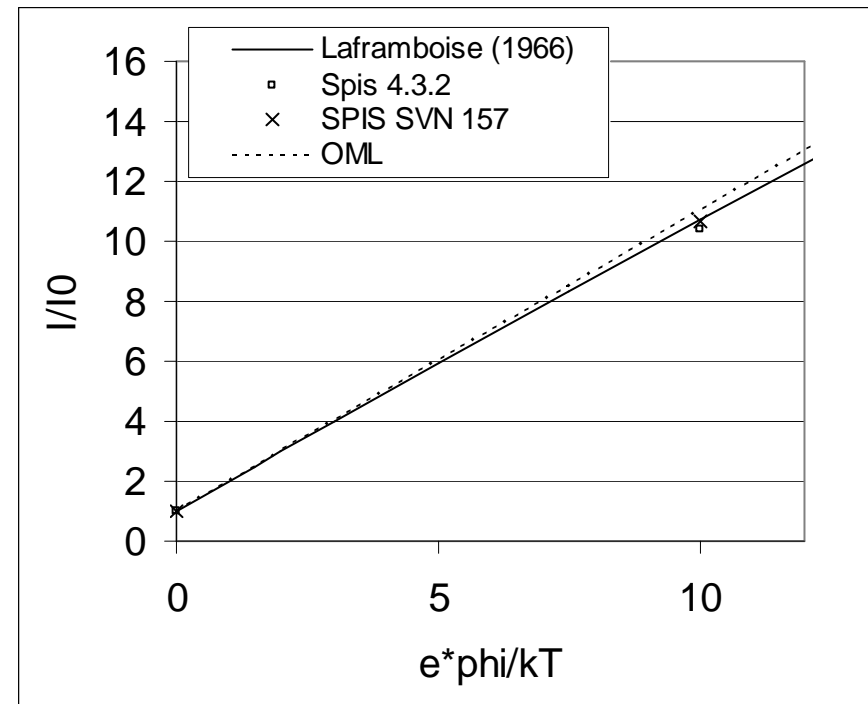
Distribution functions

- Only bi-Maxwellian environment in Spis 4.3.1
- Work progress : Generic distribution functions
 - Generic 3D sampler and isotropic samplers
 - **User-defined df : tabulated files**
 - Analytical laws can also be implemented
- Work to be performed
 - Implement analytical df: kappa, drifting maxwellian for ambient
 - Implement user-defined isotropic df for secondary electron emission and photoemission process

Boundary conditions

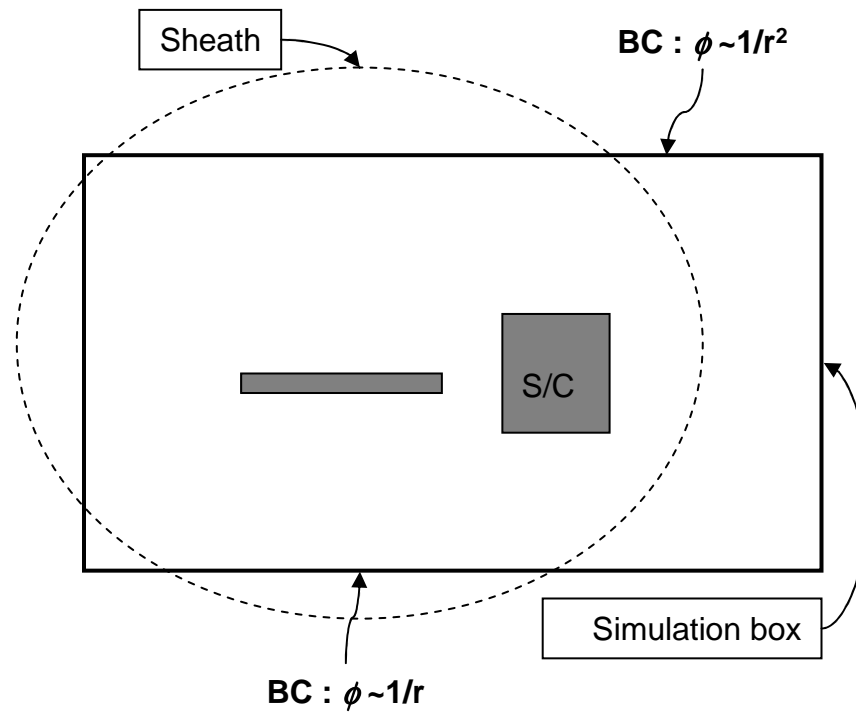
- Upgrade of particle injection
 - Non-zero electrical potential at external boundaries → disturbed plasma → change of the injected distribution function
 - Validation: better accuracy in the case of a spherical probe immersed in maxwellian plasma (error<1% instead of 3-5 % in previous version))

Quantity	Value
Temperature	0.5 eV
Electron/ion density	$6,91 \times 10^8 \text{ \#/m}^3$
Debye length	0.2 m
Potential	[0-5 V]
Sphere radius	0.1 m
Number of tetrahedrons	~42,000
Simulation box diameter	1.0 m
Number of macro-particles	~500,000



Boundary conditions

- Electrical potential
 - Automatic and local shift between two regimes
 - If the external box is outside the sheath : pre-sheath model $\phi \sim 1/r^2$
 - Else : vacuum-like condition $\phi \sim 1/r$



V cross B field

- Magnetically induced electric field

- Spacecraft motion in a magnetic field \rightarrow change of electric field $\mathbf{E}' = \mathbf{E} + \mathbf{V}_{SC} \otimes \mathbf{B}$

- Chosen implementation : mix of the 2 referential frames

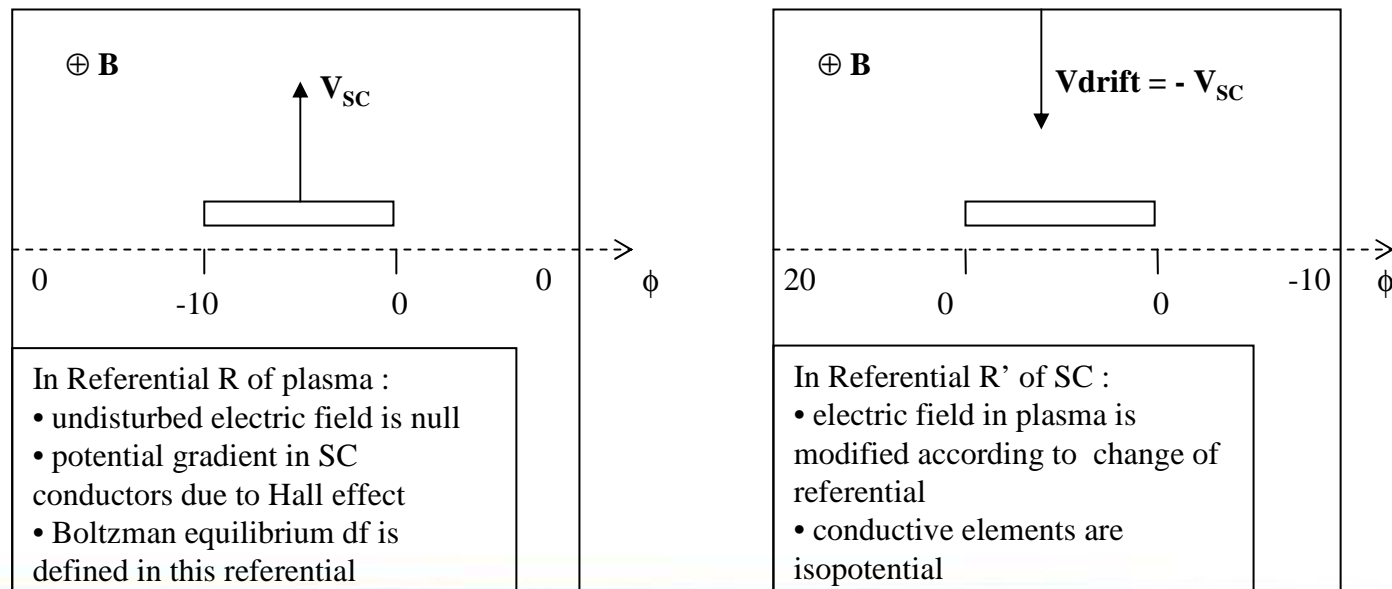
- Poisson equation in R of plasma

- BC on spacecraft due to Hall effect $\phi = \phi' + (\mathbf{V}_{SC} \otimes \mathbf{B}) \cdot \mathbf{x}$

- Particle dynamics in R' of spacecraft

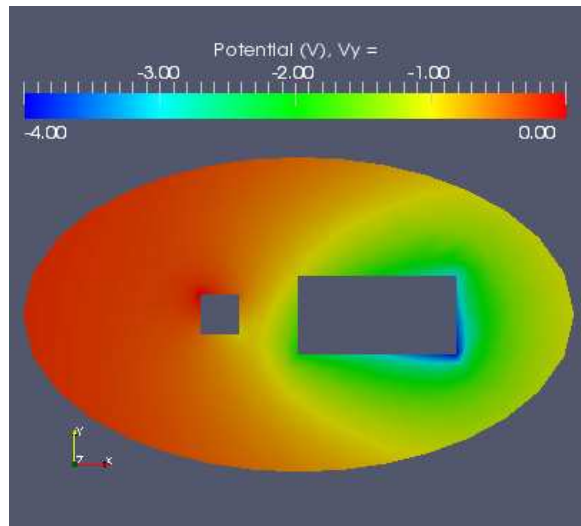
$$\dot{\mathbf{x}}' = \mathbf{v}'$$

$$\dot{\mathbf{v}}' = \frac{q}{m} (\mathbf{E} + \mathbf{v}' \otimes \mathbf{B} + \mathbf{V}_{SC} \otimes \mathbf{B})$$



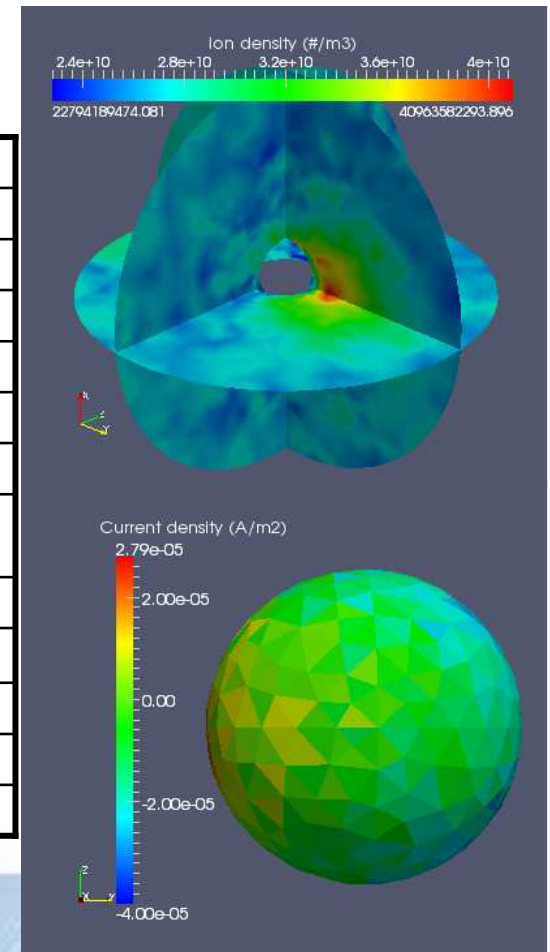
V cross B field

- Illustrative example
 - Spacecraft in LEO
 - $B \sim 5 \times 10^{-5}$ T, $V \sim 7500$ m/s, $L \sim 10$ m
 - Potential drop of ~ 4 V



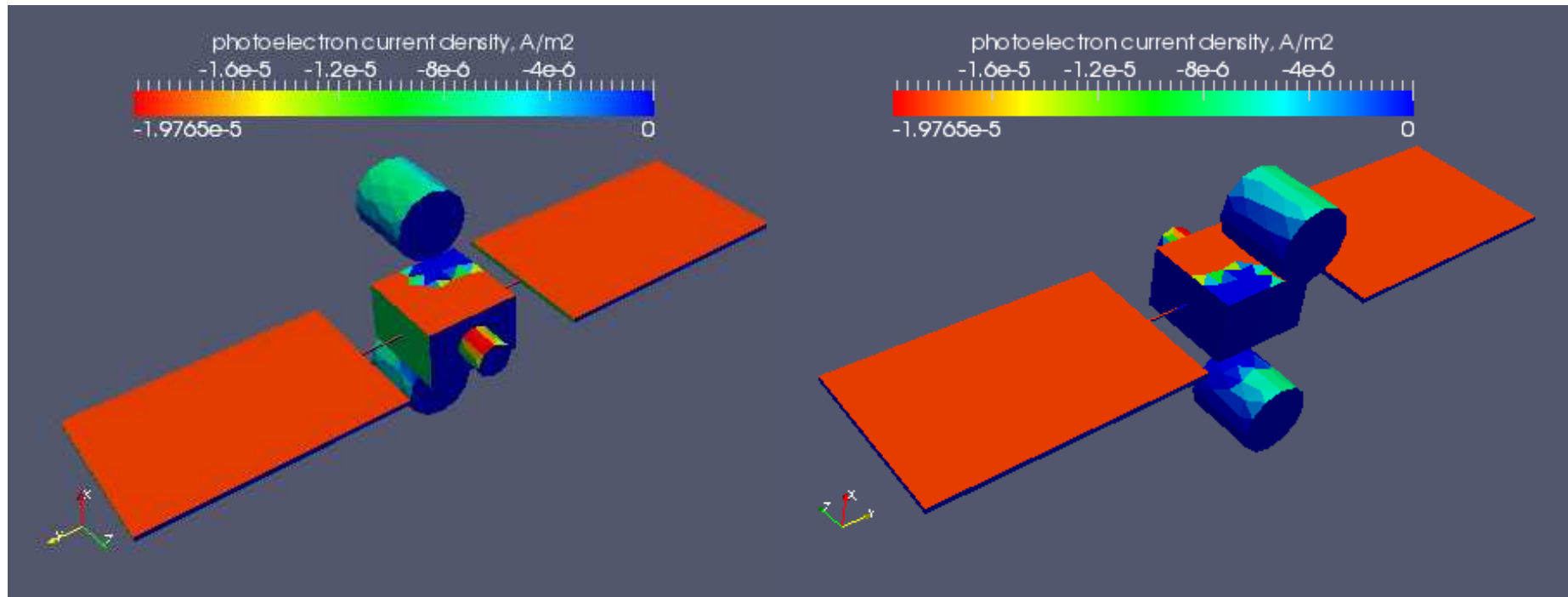
- Verification case
 - Correctly fits PTetra code: Marchand, IEEE Trans. Plasma Sci. 2012 (solved in R' of SC)

Quantity	Value
Temperature	0.2 eV
Electron/ion density	$2,8 \times 10^{10}$ #/m ³
Debye length	0.02 m
Sphere radius R	0.02 m
Number of tetrahedrons	$\sim 40,000$
Simulation box diameter	0.20 m
Number of macro-particles	$\sim 500,000$
B_z	3×10^{-4} T
Electron Gyro Radius / R	0.18
Ion Gyro Radius / R	7.6
Spacecraft velocity V_y	- 7500 m/s
Floating potential	$\sim - 0.4$ V



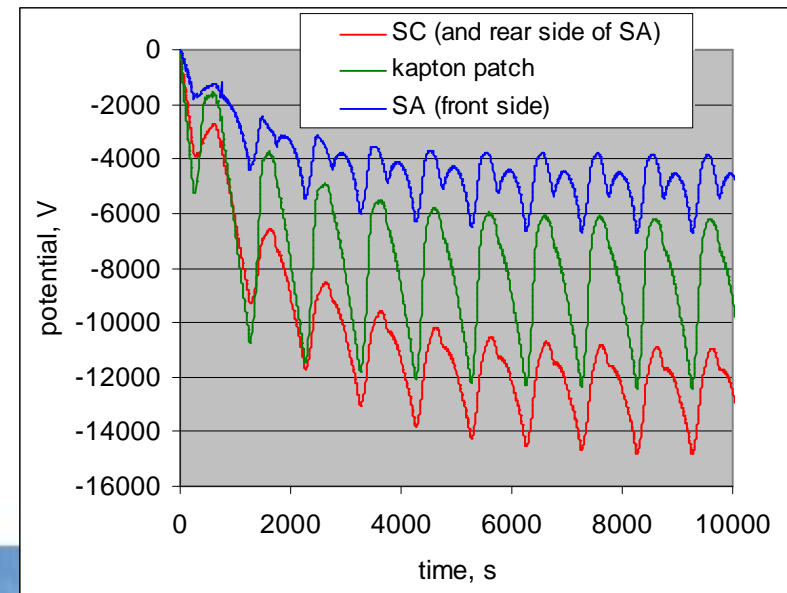
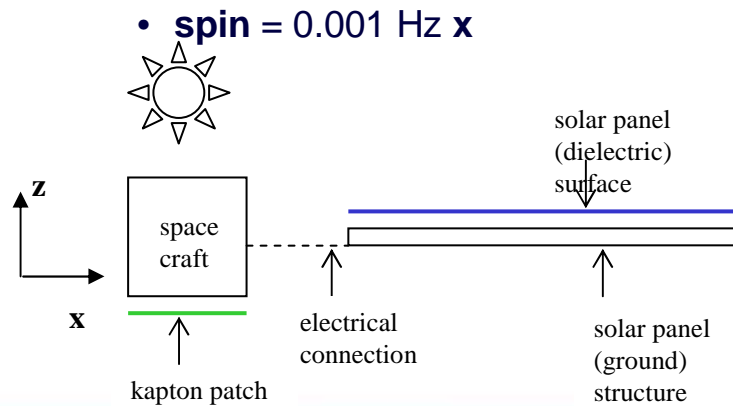
Self-shading

- In conjunction with SPIS-GEO



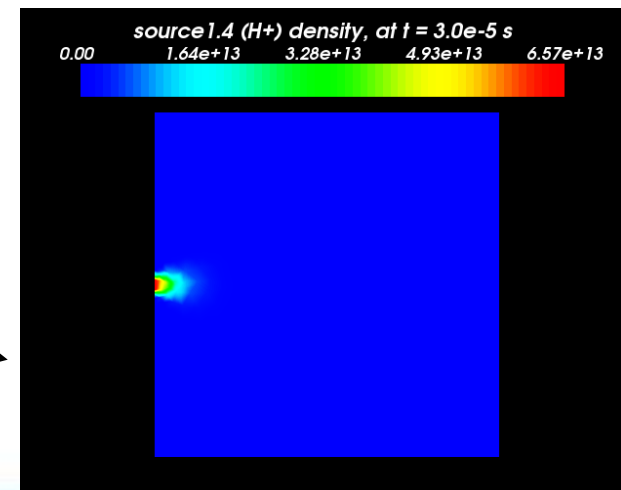
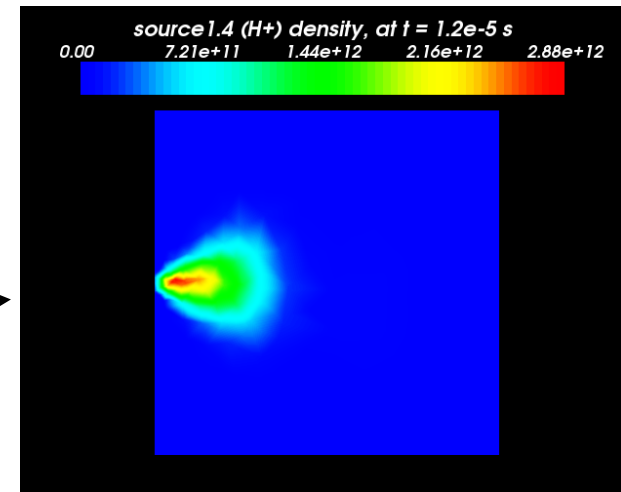
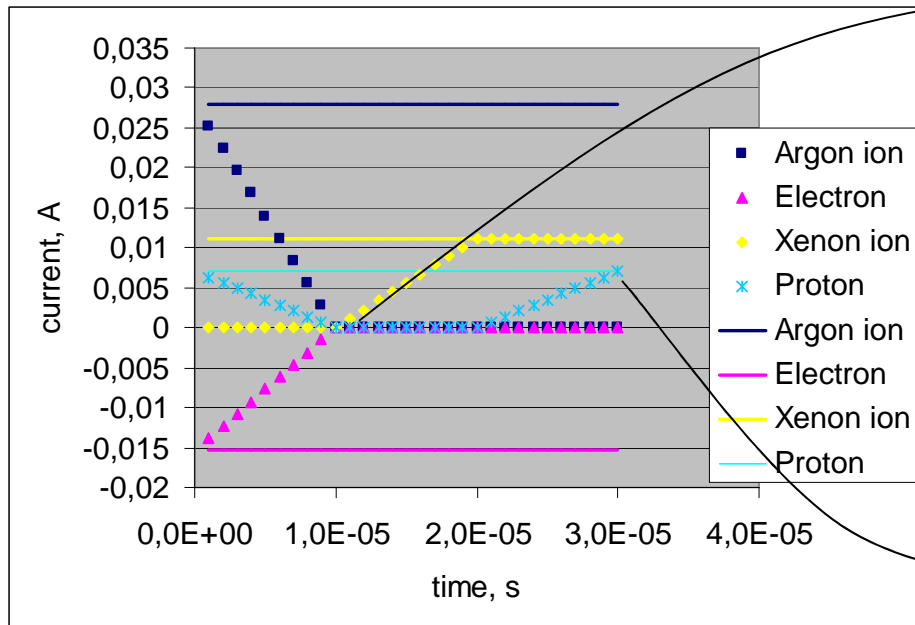
Transitions

- As of today: simulation with constant parameters
- New SPIS capability : Transitions
 - Parameters modified within the course of the simulation
 - User defined transition from
 - Global parameters for generic inputs (e.g. number, type of transitions)
 - ASCII files for dedicated inputs (e.g. sun flux change for an exit from eclipse)
- Spinning spacecraft
 - Change of sun flux (implemented) and plasma injection (not impl. yet)
 - Inputs: spin axis and angular velocity
 - Academic example of a spacecraft in GEO



Transitions

- Eclipse exit
 - see SPIS-GEO presentation
- Plasma source activation



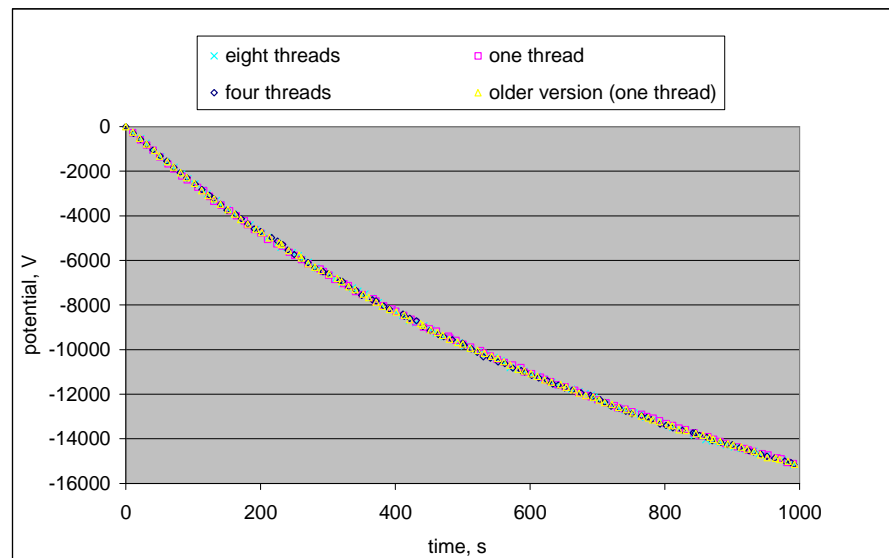
Monitoring CPU time

- Information on computational cost of SPIS numerical solvers
 - During the simulation
 - Summary at the end:

```
|-- End of SPIS numerical simulation --|
|-- Task durations
    |-- Task Simulation integration      | Cumulative duration :      25 MINUTES
    |-- Task Plasma      | Cumulative duration :      24 MINUTES
    |-- Task Plasma/SC Interactions      | Cumulative duration :       5 SECONDS
    |-- Task SC Circuit      | Cumulative duration :      46 SECONDS
    |-- Task Results storing      | Cumulative duration :       4 SECONDS
|--
|----- Plasma subtasks
    |-- Task Poisson Solver      | Cumulative duration :     131 SECONDS
    |-- Task Move all populations      | Cumulative duration :     22 MINUTES
        | At population level-----
    |-- Task Injection of ions1 density component      | Cumulative duration :       9 SECONDS
    |-- Task Push of ions1 density component      | Cumulative duration :       5 MINUTES
    |-- Task Move of ions1 density component      | Cumulative duration :       5 MINUTES
    |-- Task Injection of ions1 current component      | Cumulative duration :     10 SECONDS
    |-- Task Push of ions1 current component      | Cumulative duration :    120 SECONDS
    |-- Task Move of ions1 current component      | Cumulative duration :    131 SECONDS
    |-- Task Injection of ions2 density component      | Cumulative duration :     11 SECONDS
    |-- Task Push of ions2 density component      | Cumulative duration :    228 SECONDS
    |-- Task Move of ions2 density component      | Cumulative duration :    248 SECONDS
    |-- Task Injection of ions2 current component      | Cumulative duration :     10 SECONDS
    |-- Task Push of ions2 current component      | Cumulative duration :    100 SECONDS
    |-- Task Move of ions2 current component      | Cumulative duration :    110 SECONDS
    |-- Task Injection of photoElec      | Cumulative duration :     44 SECONDS
    |-- Task Push of photoElec      | Cumulative duration :    210 SECONDS
    |-- Task Move of photoElec      | Cumulative duration :    261 SECONDS|
    |-- Task Injection of secondElec      | Cumulative duration :     40 SECONDS
    |-- Task Push of secondElec      | Cumulative duration :    223 SECONDS
    |-- Task Move of secondElec      | Cumulative duration :    268 SECONDS
|-----
```

Multi-threading "parallelisation"

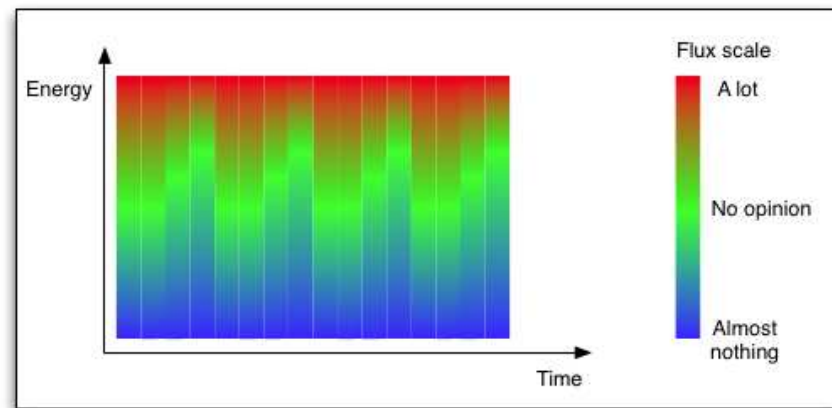
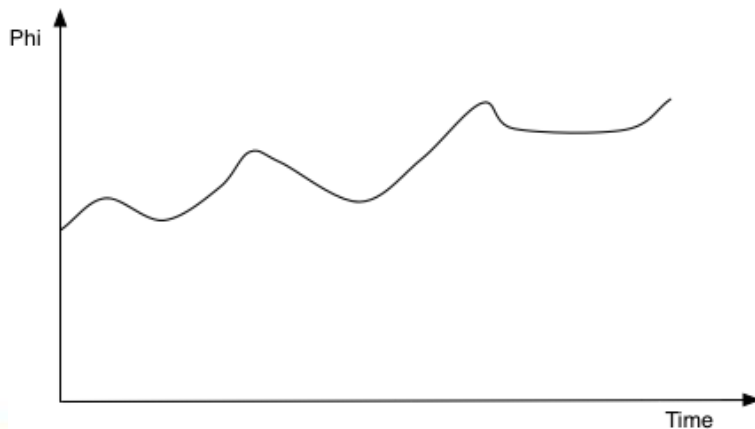
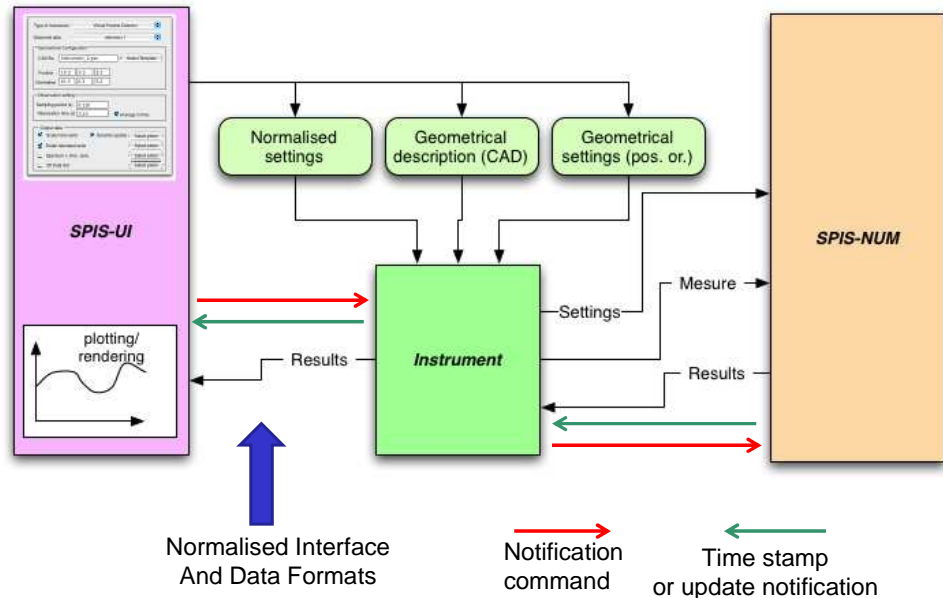
- The particle pusher is often the main CPU time consuming calculation
- Multi-threading approach
 - Based on Java Thread class
 - Good scale (ThreadNb vs. CPU time) when the particle pusher is very costly (lots of particle and large integration times)
 - Permit to reduce the cost of particle pushing to the same level as Poisson equation



Scientific Instruments Design

- Plasma Sensors and Particle Detectors

- Location
 - On single point of the plasma volume
 - On an instrument (possibly virtual)
- Temporal measurement
 - Regular observations
 - On user demand (interactive mode)
- Outputs
 - Distribution functions and first momentum
 - Plasma potential



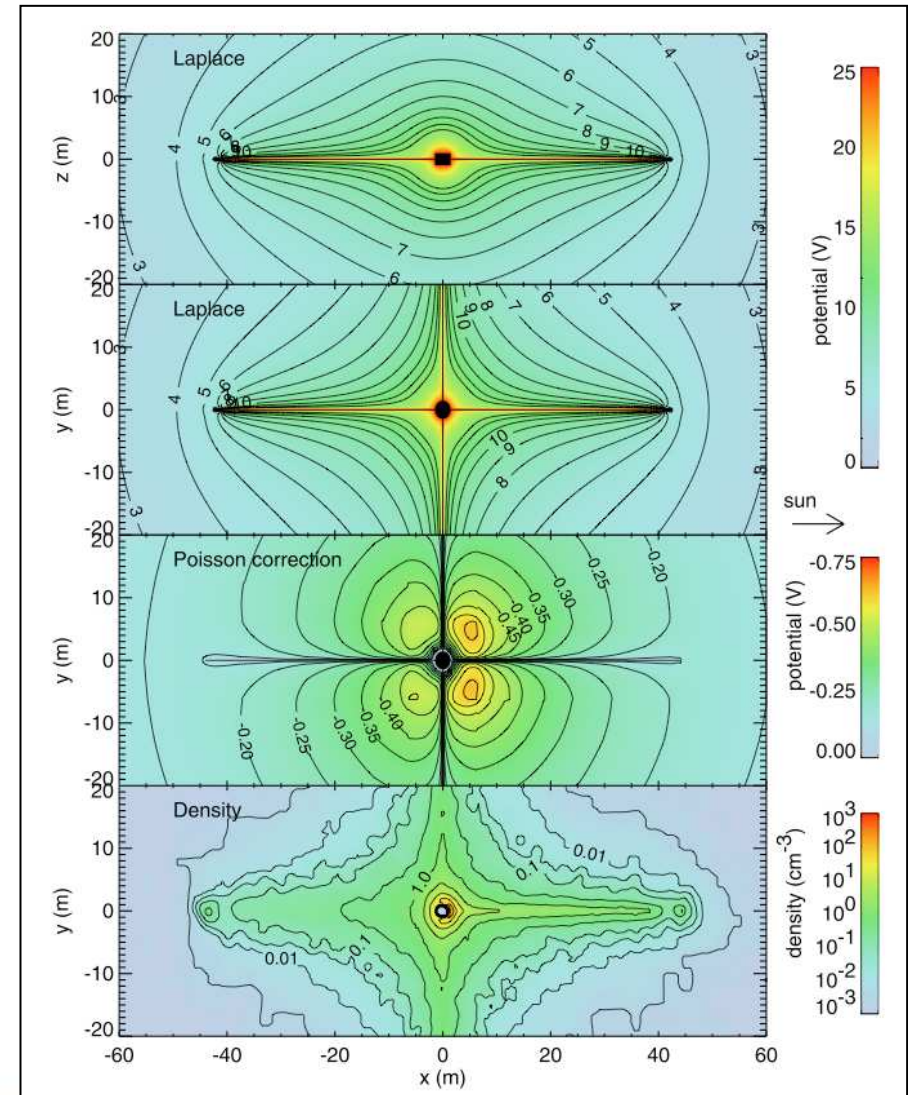
Scientific missions modeling – Validation cases

- 5 validation cases / 3 missions have been defined
 - Currently under development by specialists of electric field sensor and particle detectors involved in several scientific missions : IRF (Sw) and IRAP (F)

Case	1	2	3	4	5
Scope	Cluster E-field measurements	Cluster electron measurements	Solar Orbiter E-field measurements and wake	Solar Orbiter electron measurements	Cassini electron measurements
S/c relevance	Bepi MMO MMS	Bepi MMO MMS	Rosetta JGO SP+ Demeter	Solo Bepi MPO Rosetta SP+	JGO Rosetta
Plasma conditions relevance	Bepi MMO MMS	Bepi MMO MMS	Bepi MMO Bepi MPO SP+	Bepi MMO Bepi MPO SP+	JGO Rosetta Demeter
Relevant existing databases	Cluster	Cluster	Rosetta	Rosetta Helios Stereo	Cassini
Relevant previous studies	[Cully2007] [Mihaljic2010] [Miyake2011] [Thiébaud2003] [Thiébaud2004]	[Mihaljic2010] [Pedersen2008] [Szita2001] [Thiébaud2003] [Thiébaud2004]	[Engwall2006]	[Ergun2010] [Isensee1981] [Katz2001] [Sjögren2009]	[Jacobsen2009] [Laframboise1974] [Lewis2010] [Nilsson2009] [Olson2010]

Validation Case 1 - Cluster E-field

- Cluster attractive features
 - Complete set of high-sensitivity plasma instruments onboard
 - Large and accessible database
 - Several studies on s/c-plasma interaction
 - High interest from the scientific community
 - Relevance for Bepi-Colombo MMO, MMS, Themis ...
- Model
 - Spacecraft, wire booms, spherical probes and adjacent electrical elements
 - From 0.3 mm thickness to 88 m probe-to-probe distance
 - Currents to small bodies crucially depends on test particle tracking
- Comparison
 - Other codes (Miyake2011, [Cully-2007](#))
 - In-orbit study (Mihaljic2010,)

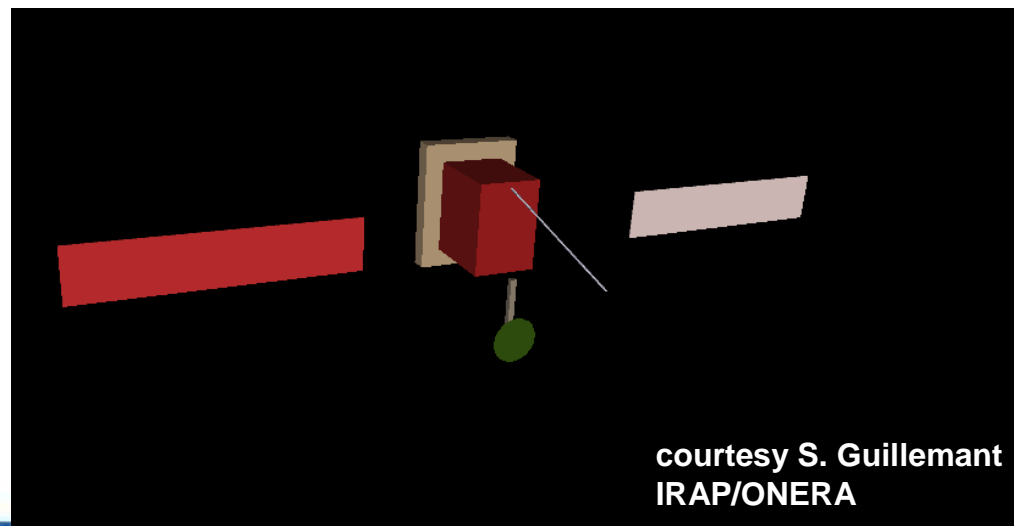


Validation Case 2 - Cluster electron measurement

- Different conditions
 - Tenuous Solar wind at Earth orbit
 - Dense Solar wind at Earth orbit
 - Polar cap
- Model
 - Electron detectors
- Objectives
 - Quantification of photoemission and secondary emission
- Comparison
 - measurement with PEACE

Validation Case 3 - Solar Orbiter electron measurements

- Objectives
 - Electron instrument (located on an anti-sunward boom)
 - Potential barrier can form around the spacecraft out of combined effects of photoelectron emission and ambient electron density
 - Quantification of the photoemission and of the secondary emission
 - Effect of \neq Distribution functions: Maxwellian, Kappa or user defined from observations
- Comparison
 - SPP simulations [Ergun2011]
 - Distribution functions obtained by a simulated electron instrument (SWA/EAS).

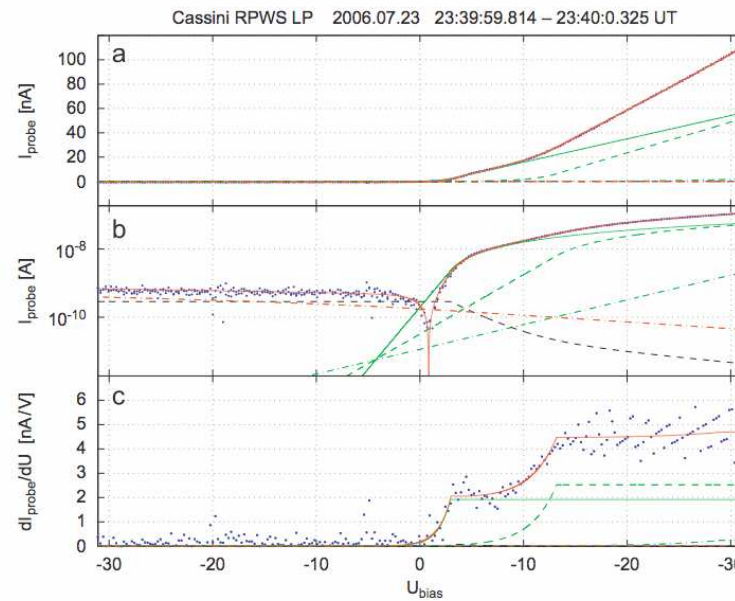
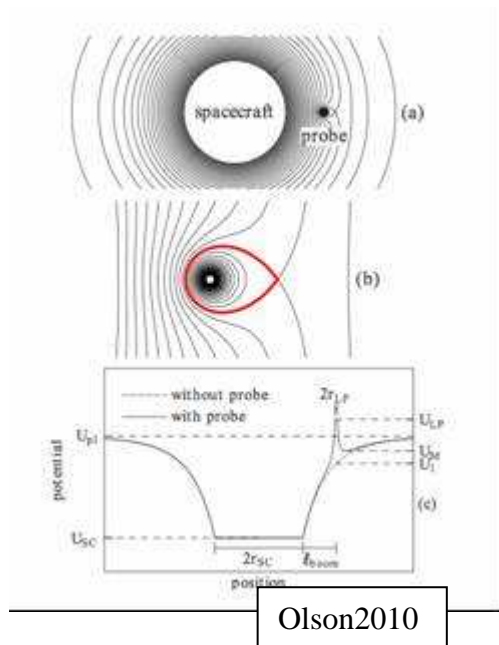


Validation Case 4 - Solar Orbiter E-field and wake

- Objectives
 - Wake effect on the electron instrument
 - Electric potential pattern near the antennas
 - Possibly several solar panel orientations
 - Effect of non-conducting solar panels on electric field measurements

Validation Case 5 - Cassini

- Objectives
 - Current-voltage characteristics of the Langmuir probe



Validation Case 5 - Cassini

- Other objectives still to be defined in detail
 - Influence of photoelectrons and spacecraft potential on electron measurements of RPWS-LP and ELS @ negative and positive spacecraft potential
 - Effect of secondary electron emission on the RPWS-LP bias voltage sweeps
 - Some sweeps [Garnier2011] show a region of negative resistance
 - Direct effect of secondary emission
 - Or Bfield effect and wake effect
- Model
 - Particle detectors and Langmuir Probes
- Comparison
 - A good set of high-sensitivity plasma instruments onboard
 - Large and accessible database, several studies on SC interaction and measurement [Jacobsen2009, Nilsson2009, Olson2010, Lewis2011]

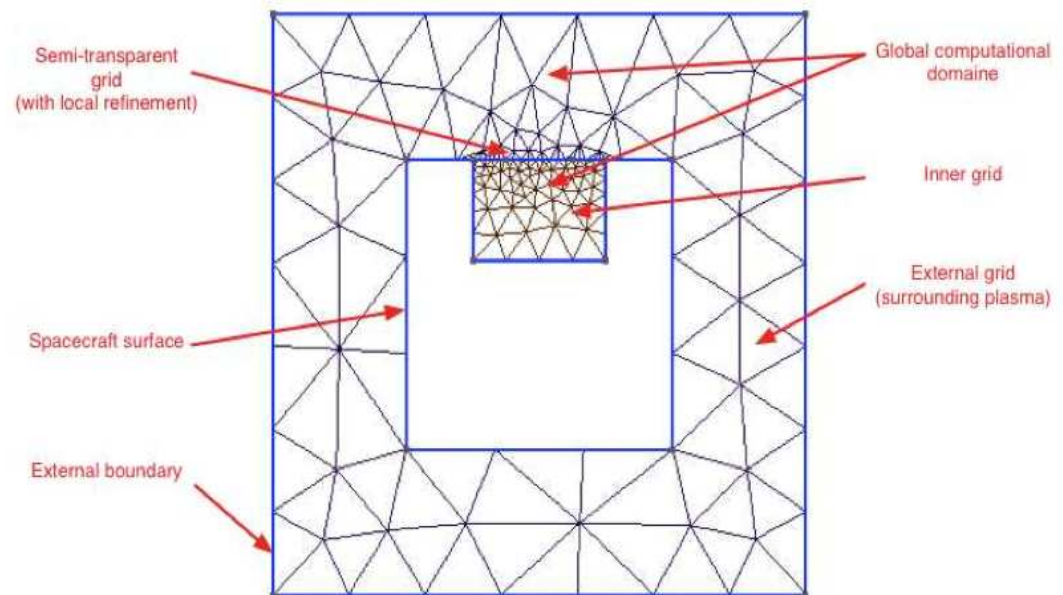
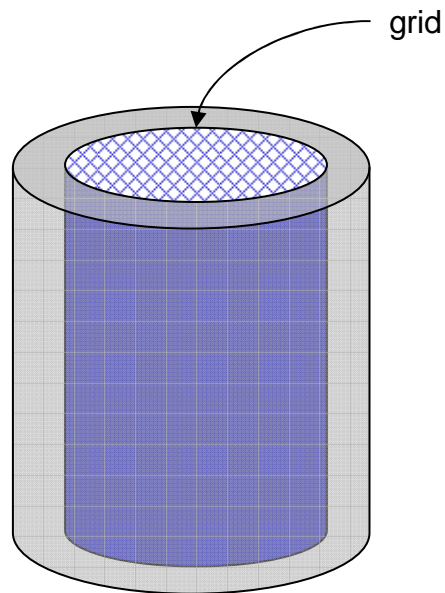
Summary

- Thanks SPINE community for providing ideas of SPIS development and validation cases !
- See you at 12th SCTC and next SPINE WS !

Reserves

Semi-Transparent Grids

- Useful to model some specific instruments
- Work performed
 - Definition of the STG format: mesh, dependencies of the circuit solver, particle pusher



Particle Detectors Design

Key aspects

- provide good statistics results on small instruments compared to SC

Test Particle (TP) method

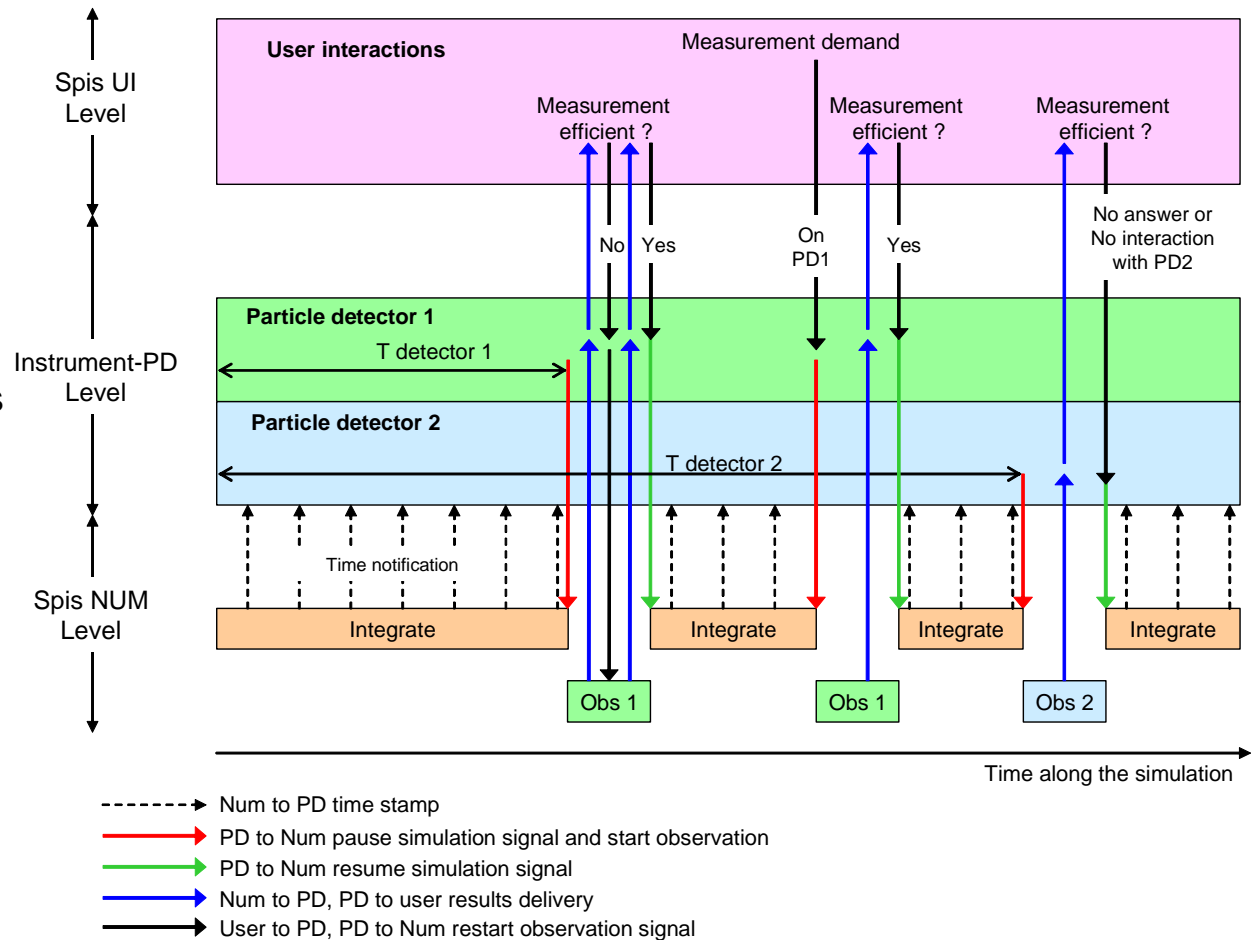
- “Frozen” simulation
- Mix of Forward tracking and Backtracking of TP populations
- Increase statistics on instruments surfaces

Settings

- Pre-defined parameters
- Interaction with the user
 - modify energy range
 - super-particle number

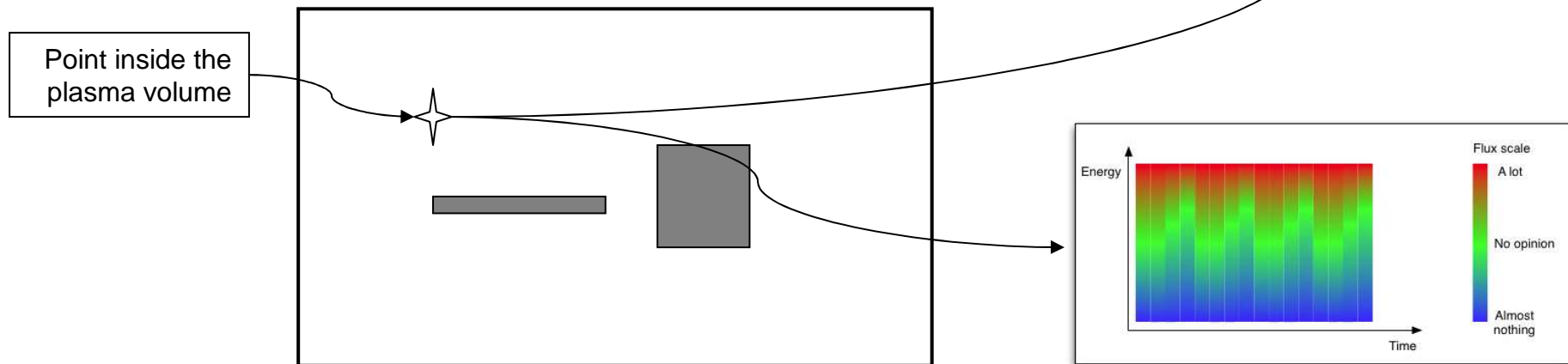
Work progress

- Design of new package
- UI / Instruments / NUM package interfacing



Plasma Sensors progress

- Provide light data concerning the simulation
 - Plasma potential
 - Energy distribution function and density
 - At a single point of the simulation
 - Using the PIC populations of the simulation (no TP)



- Useful to check the convergence of the simulation
- Work performed so far
 - Numerical routines OK
 - User Interface wizards : still to be implemented