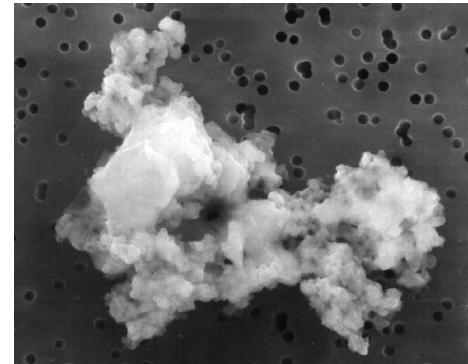
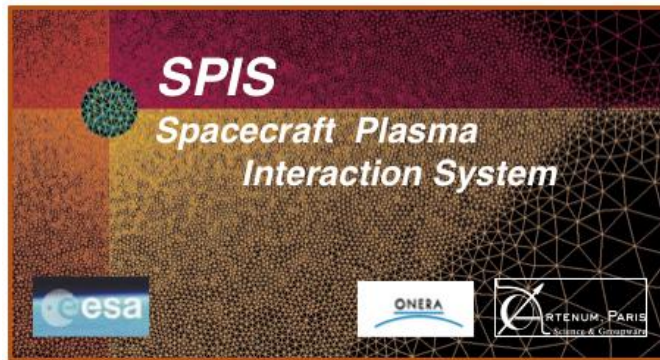

SPIS-Dust Validation Cases

19 May 2015

F. Honary, S.R. Marple, S.L.G. Hess, P. Sarrailh, J.-C. Mateo-Velez, B. Jeanty-Ruard, A. Anuar, F. Cipriani, A. Hilgers, D. Rodgers, J. Forest and B. Thiebault.

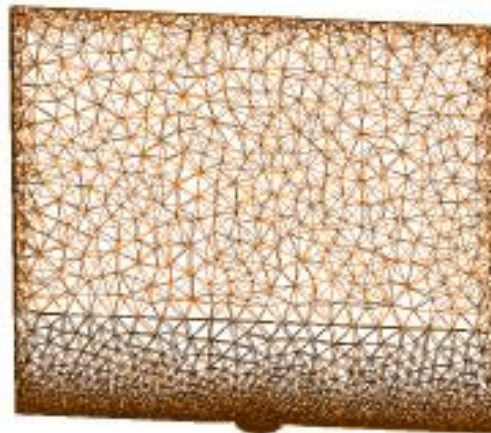
SPIS-Dust Validation Tests



2D and 3D simulation cases

2D simulation Domain

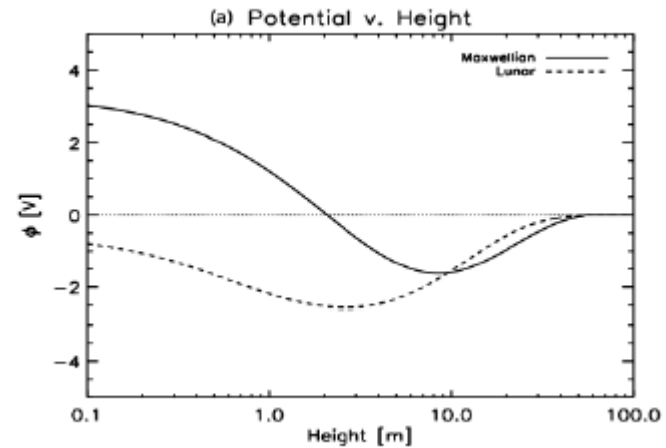
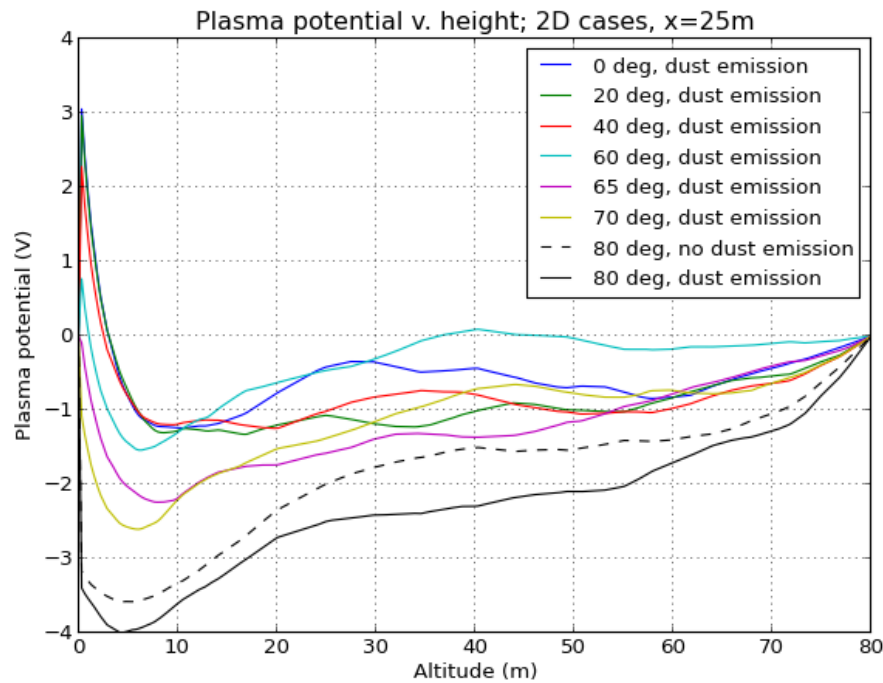
X=-50 to 50m
Y=0 to 2m
Z=-2 to 80m



Top, open
boundary, 0V
5m mesh

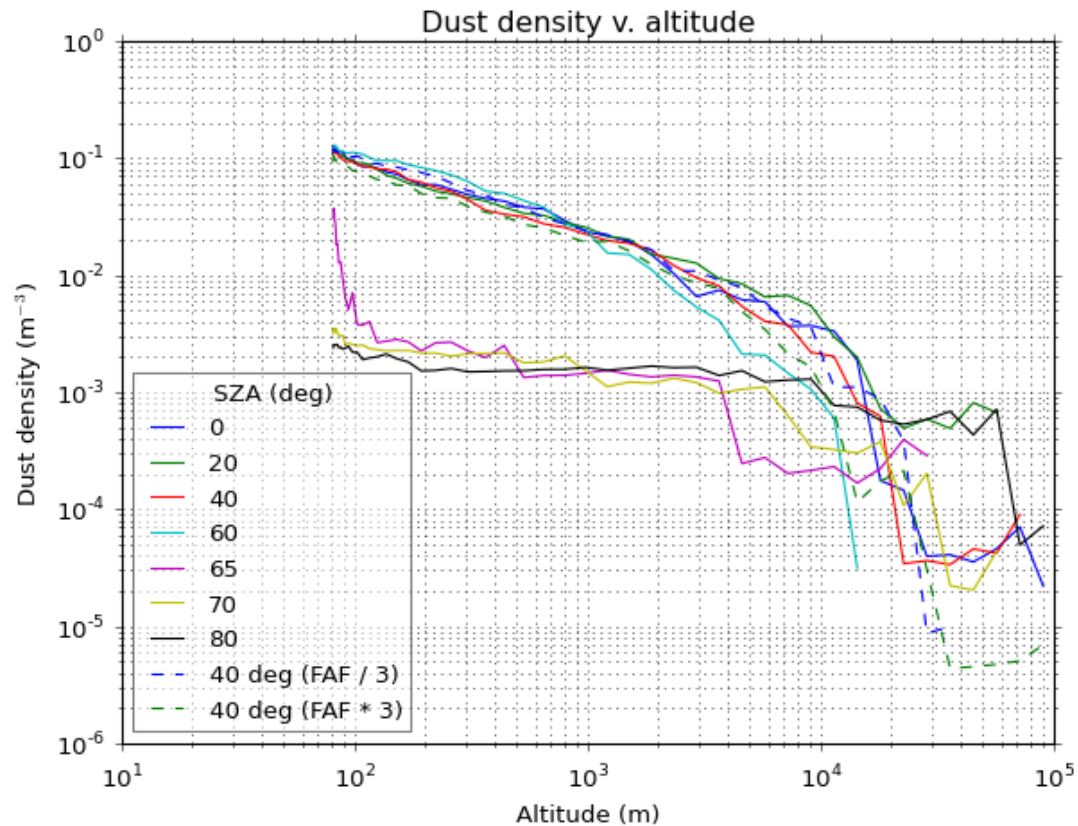
Crater, non-conductive
(dielectric) lunar dust
10m diameter
2m deep, 0.5m rim

Plasma Potential

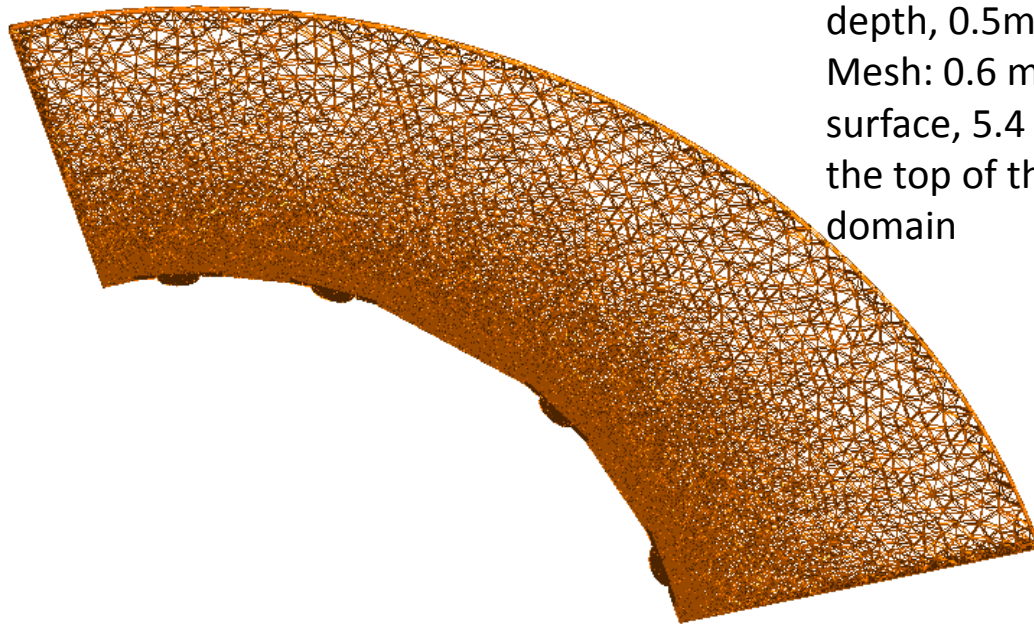


Poppe and Horanyi 2010

Dust Density for different SZA 2D simulation



Simulation Domain for Asteroid



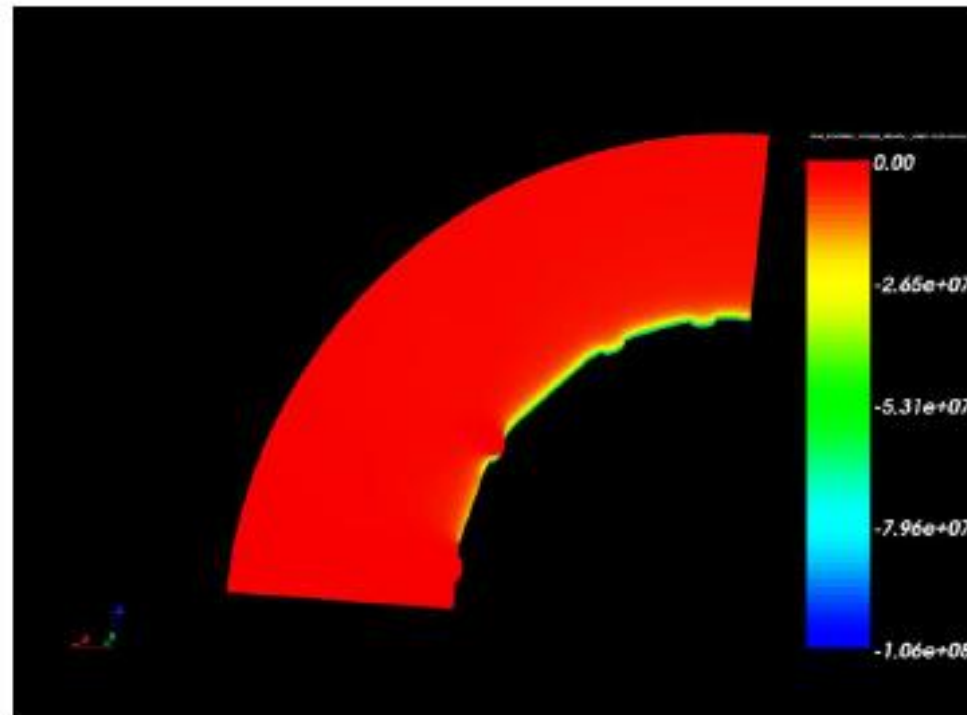
Craters: 10m
Diameter, 2m
depth, 0.5m rim
Mesh: 0.6 m at the
surface, 5.4 m at
the top of the
domain



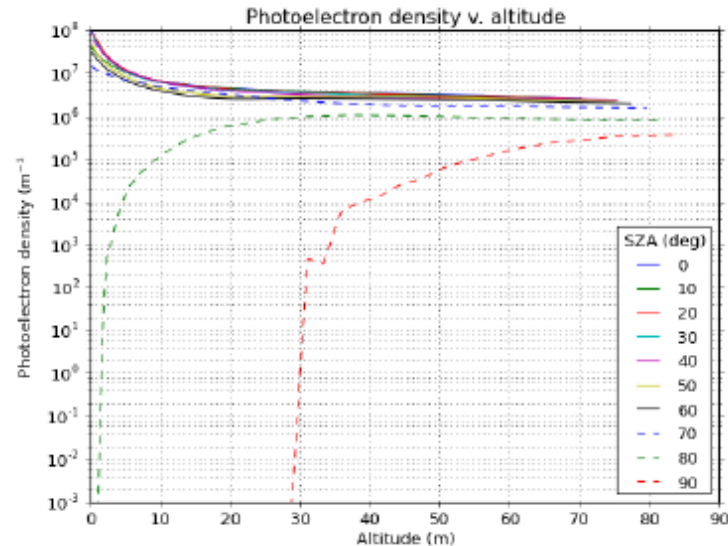
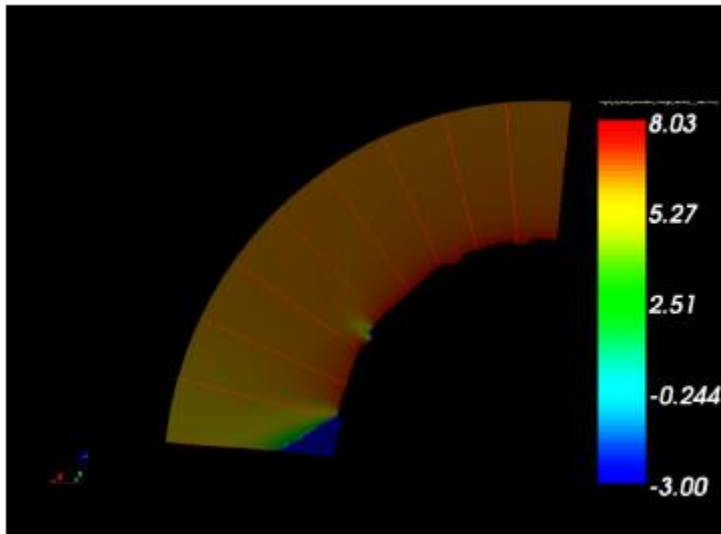
Simulation Parameters

Parameters	Values
Semi-minor radius (x axis)	100m. Simulation domain extends 85m above asteroid surface along X axis.
Semi-major radius (Z axis)	115m. Simulation domain extends 70m above asteroid surface along Z axis.
Y	-1 to 1 m

Photoelectron Sheath

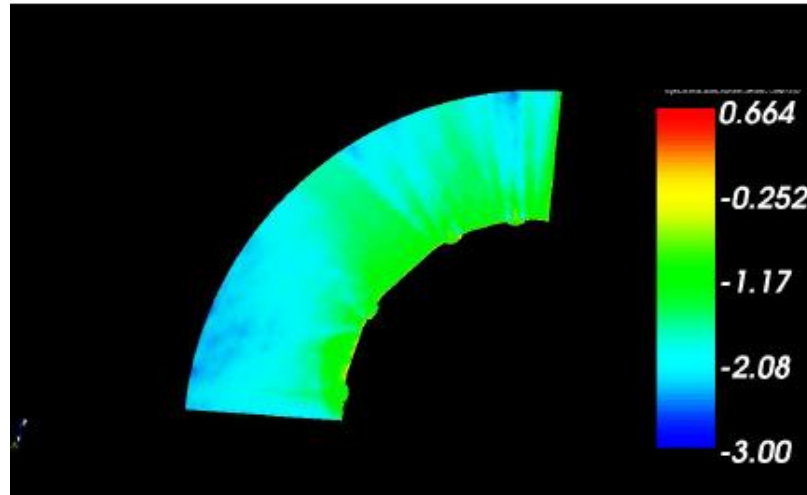


Photoelectron density as a function of altitude

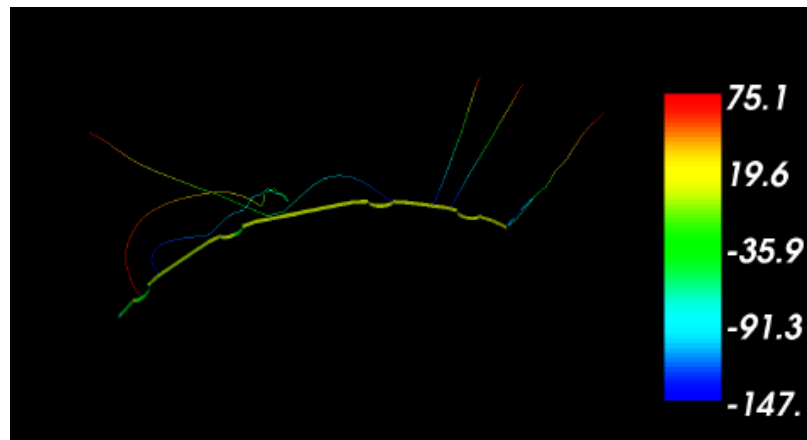


\log_{10} of photoelectron charge density ($\text{m}^{-3} \text{ecu}$) with reduced transparency to highlight the probing lines; SZA = 90° at 9 o'clock, 0° at 12 o'clock.

Dust Density and trajectory

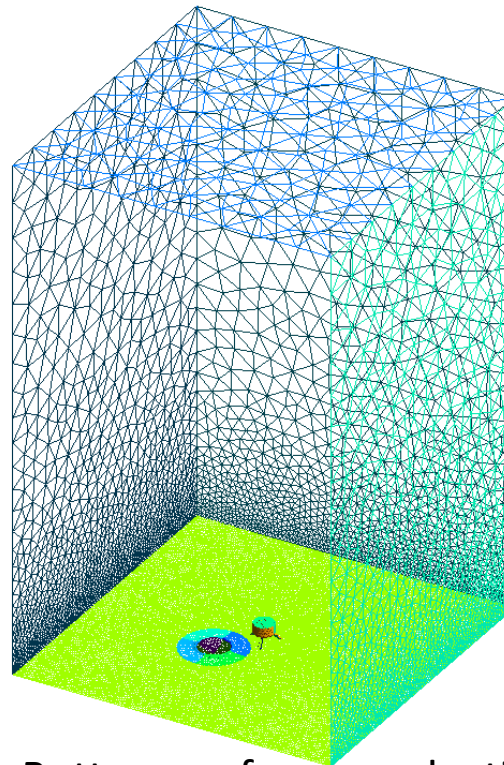


Log10 of dust density



3D Simulation Domain

Crater, non-conductive
(dielectric) lunar dust
5m diameter
1m deep



Top, open boundary, 0V
6m mesh

Sides, periodic boundary
80m high

Lander 3m diameter
Solar cells on upper part

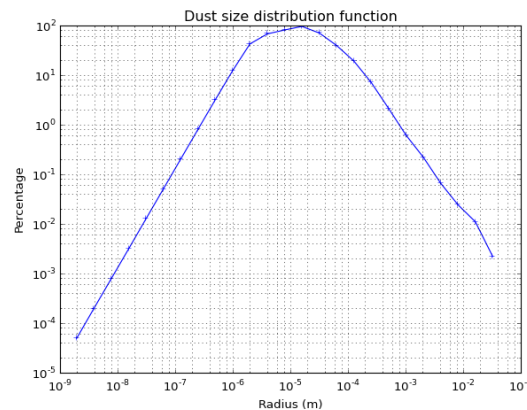
Lander underside and legs
aluminum

Bottom surface, conductive
lunar dust
50 x 50m; 0.6m mesh



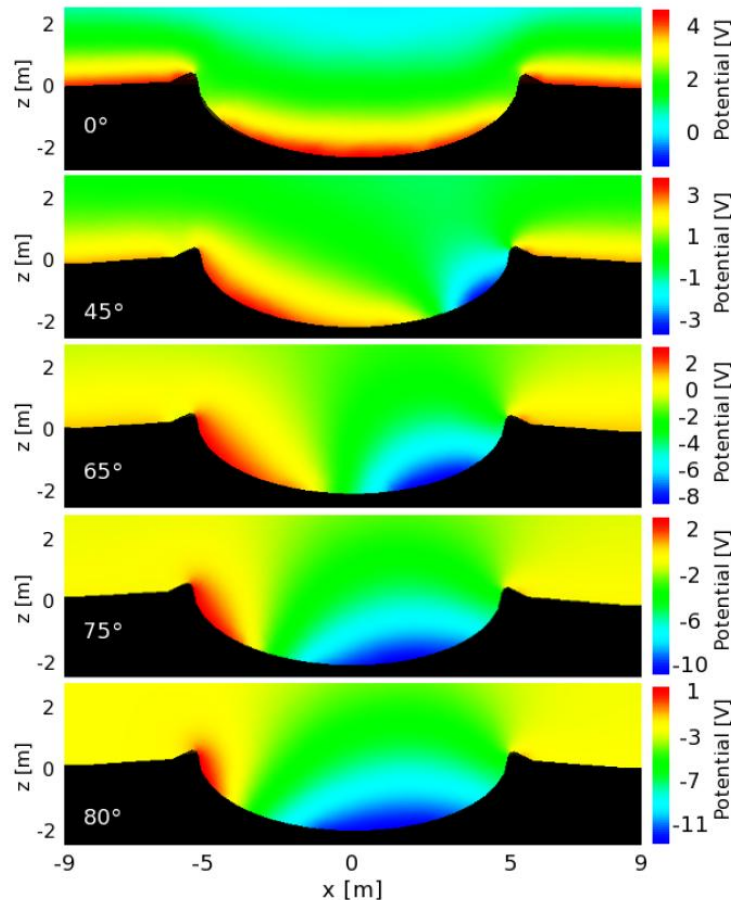
Simulation Parameters

Parameter	Value
Solar wind speed	400km/s
Electron and ion densities	$1 \times 10^7 \text{ m}^{-3}$
Electron and Ion temperatures	10 eV
Photoelectron distribution function	Maxwellian



Dust Distribution: taken from Lunar source book 1991, extrapolated for dust particles $< 1 \mu\text{m}$

Electric Potential for different SZA

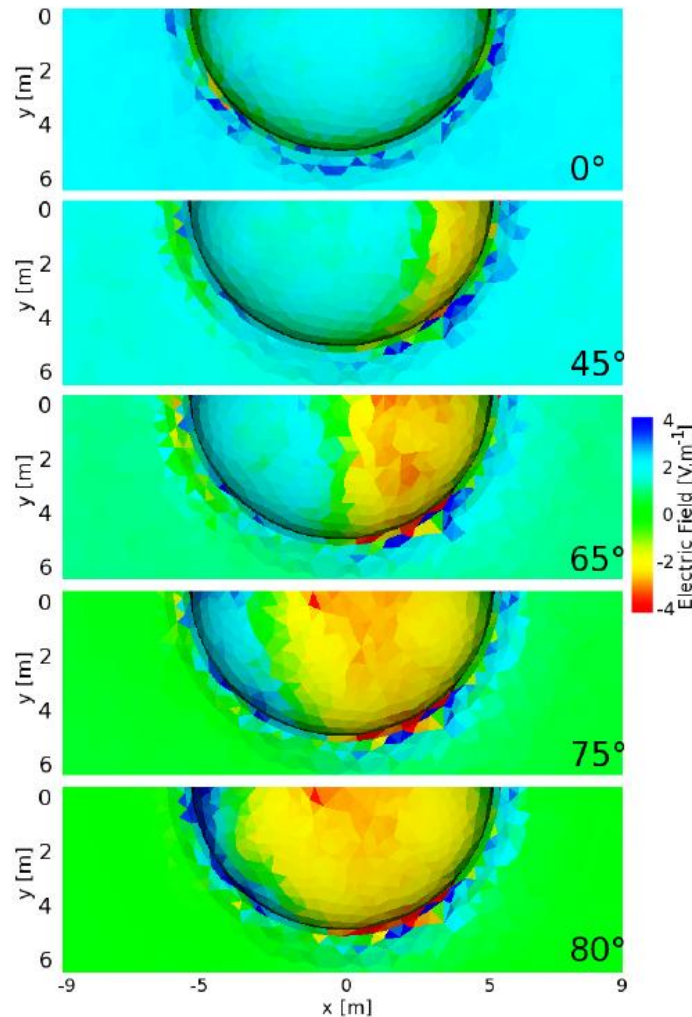


Because of both the sunlight and solar wind ion shadowing by the crater rim, the inside of the crater becomes very negative, down to potential of -15V in for SZA=82

The sunlit edge of the crater stays at a nearly constant potential of 1 - 2 Volts.

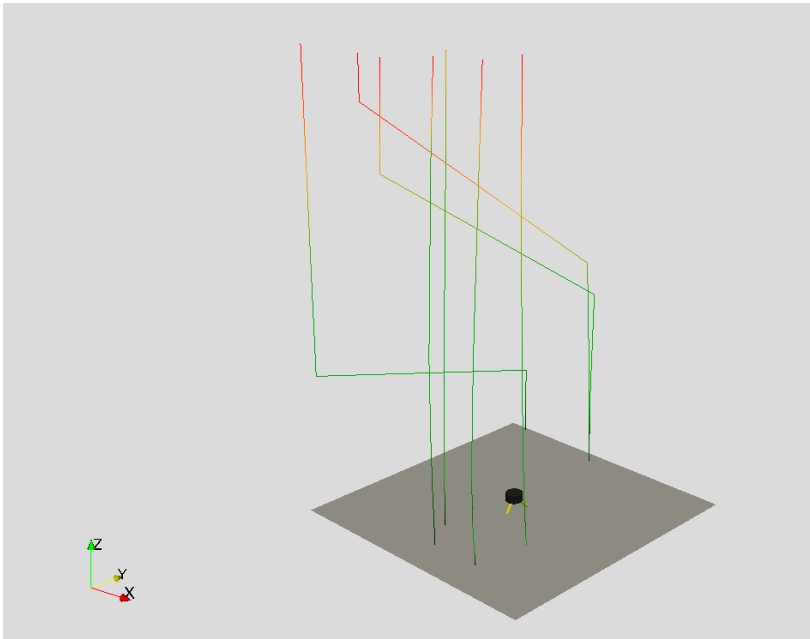
The flat surfaces out of the crater are positively charged for SZA<70 and negative for larger SZA.

Electric field normal to the surface

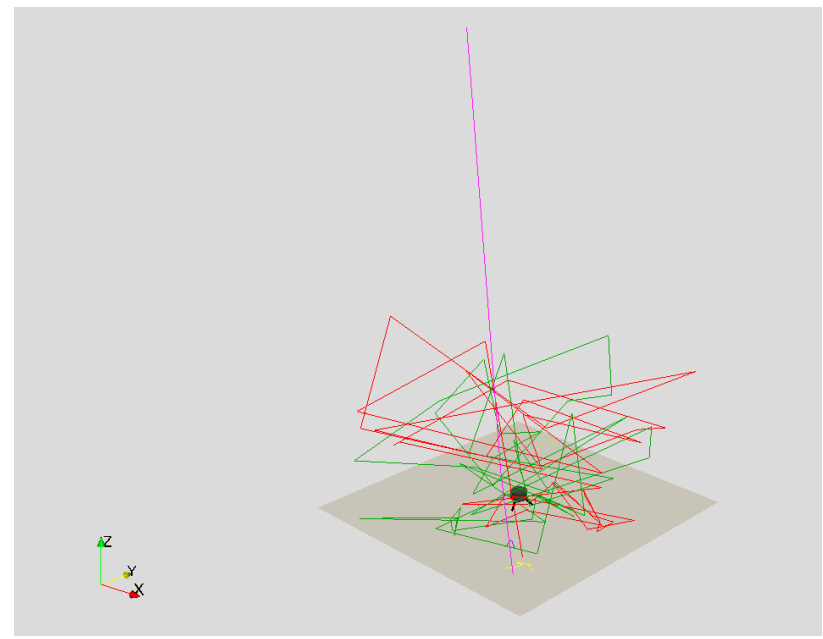


At low SZA, the electric field is mostly positive on the whole surface, leading to a positive charging of the surface dusts, whereas for large SZA the electric field is nearly zero for at surfaces, and mostly negative in the crater, leading to the negative charging of the dusts. Since the dust charge positively once emitted due to photoemission, the low SZA case seems more suitable to generate dust levitation, contrary to usual expectations.

Dust trajectory

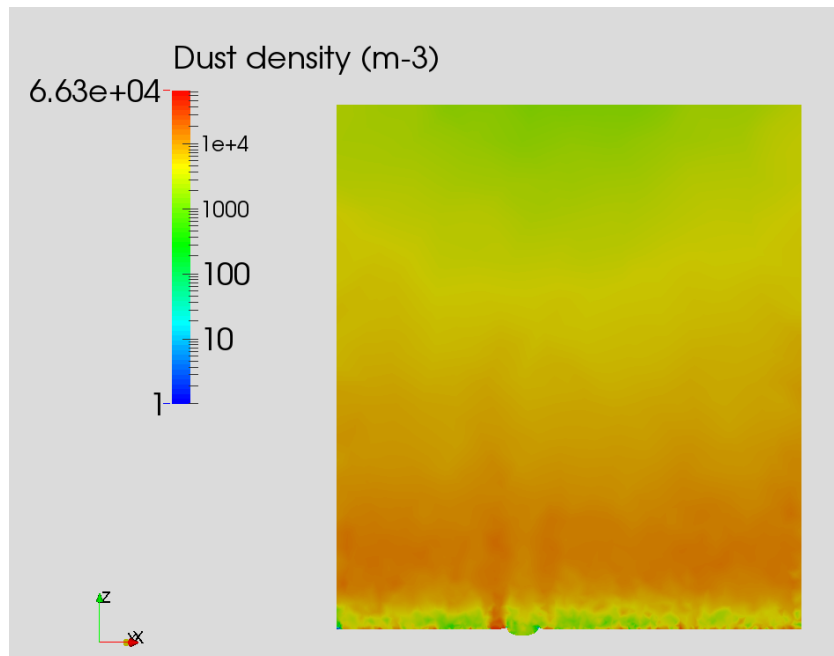


SZA=0, Lander

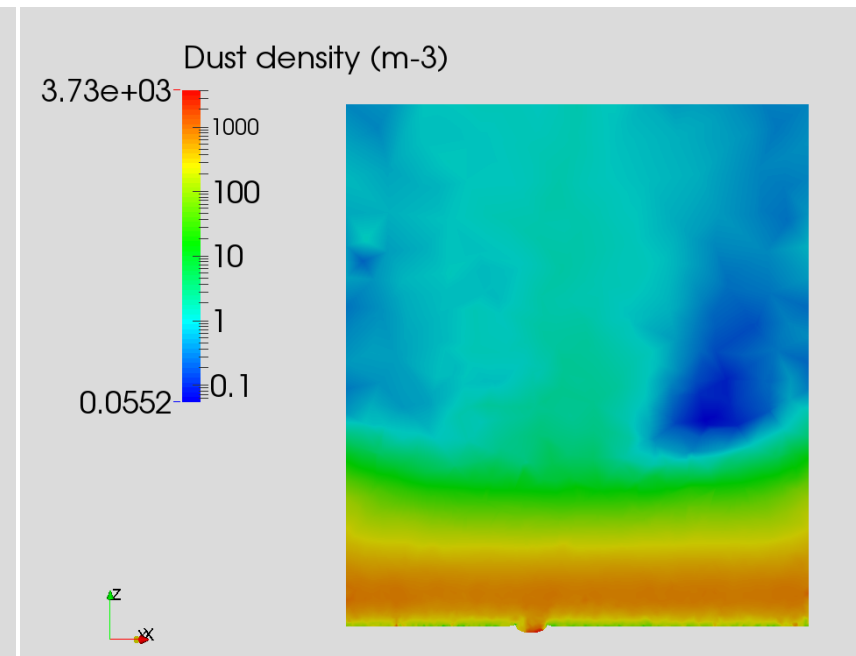


SZA=80, Lander

Dust Density in the vicinity of Crater

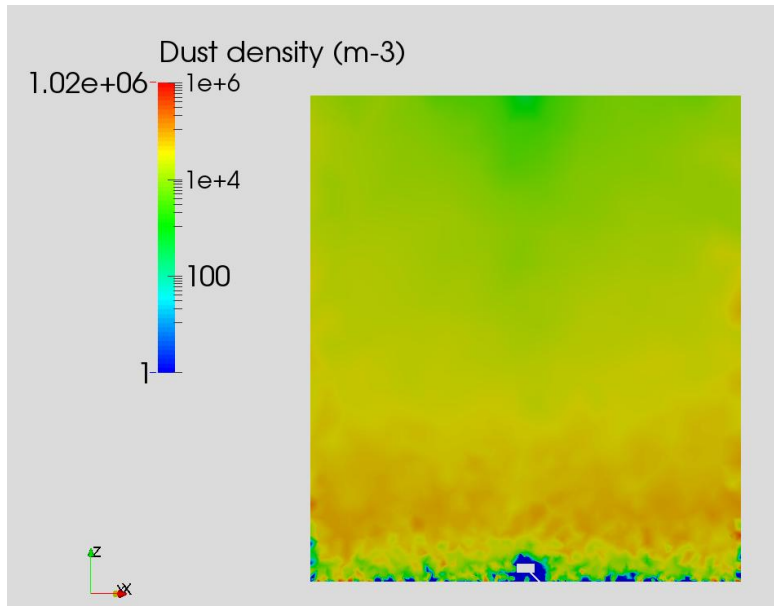


SZA=0, Crater

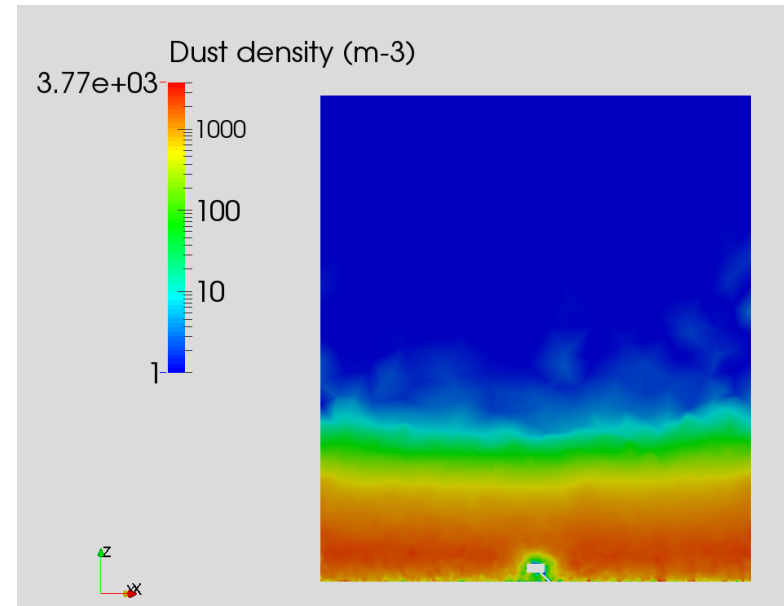


SZA=80, Crater

Dust Density in the vicinity of Lander

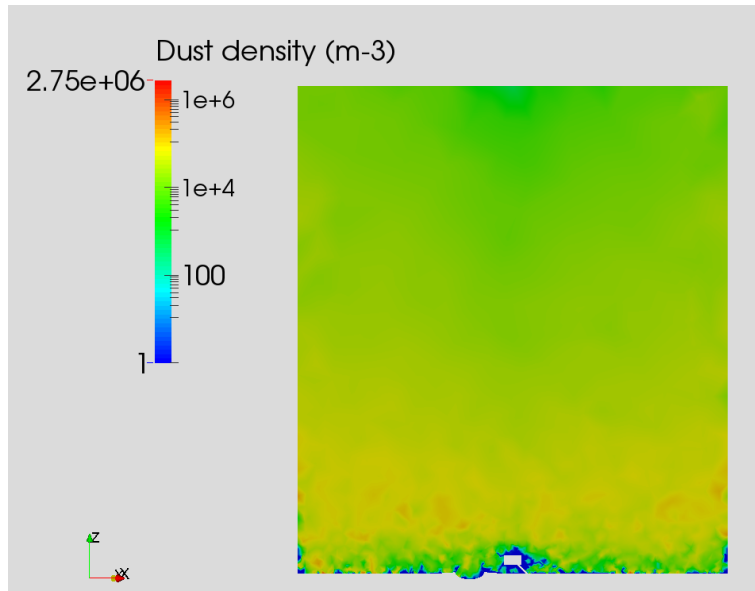


SZA=0, Lander

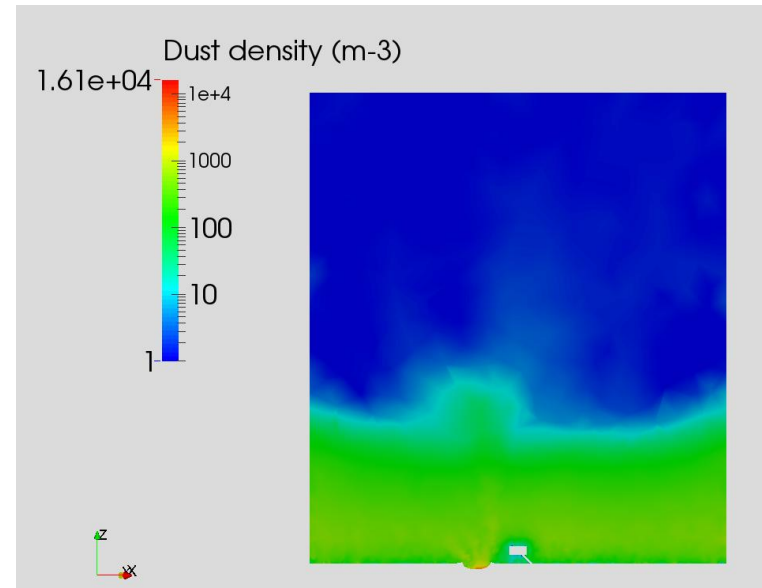


SZA=80, Lander

Dust Density in the vicinity of Crater and Lander



SZA=0, Lander and Crater



SZA=80, Lander and Crater

Thank you for listening

